National Radio Astronomy Observatory Tucson, Arizona

February 28, 1992

MEMORANDUM

To:

R. Hall, GBT Memo Series

From:

James Lamb

Subject:

Paint Performance at Millimeter Wavelengths

I. Introduction

This note presents the results of measurements of the loss at millimeter wavelengths of the paint to be used for the GBT panels. One set of measurements was performed under contract by the National Physical Laboratory, England. The other was done at NRAO. From the measurements it may be concluded that the dissipative loss in the paint is negligible up to 100 GHz, and probably to 250 GHz. There appears to be some scattering loss (of order 5%), although the effect was difficult to measure because of the lack of flatness of the samples.

II. Loss at 90 GHz

An SIS radiometer was set up and tuned to 90 GHz. An aluminum mirror was placed at 45° in front of the receiver to deflect the beam horizontally through 90° (Fig. 1). A liquid nitrogen cold load was set in the beam as close to the mirror as

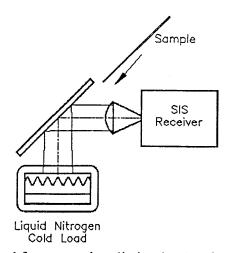


Figure 1. Setup used for measuring dissipative loss in samples at 90 GHz.

possible. When the samples were placed on the front of the mirror, the output power was constant to within 0.01 dB. Knowing that the receiver noise temperature is about

February 28, 1992 Page 2

50K, it may be deduced that the loss in the dielectric is less than 0.3% relative to the aluminum mirror.

III. Scattering at 90 GHz

To measure the scattering, the cold load was moved to about 2 m from the mirror (Fig. 2). A piece of absorber with a 6 cm diameter hole was fixed in front of

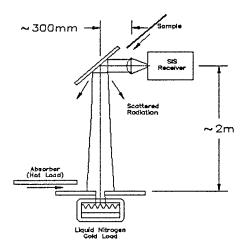


Figure 2. Setup used for estimating scattering loss at 90 GHz.

the cold load so that only the central portion of the beam was intercepted. A reading of the total power was taken, and then another with the hole covered by a piece of absorber at ambient temperature. Another set of readings was taken with each of the sample plates in front of the aluminum mirror. The ratio of the hot-cold differences then gives the loss. Since this is measured for the center of the beam and it is known from the previous section that there are no dissipative losses, it must be a direct measure of the scattering loss.

Since the plates were not perfectly flat, they were clamped to the aluminum mirror to reduce the buckling as much as possible. This was not entirely successful since there was considerable scatter in separate measurements of a single sample. In addition, some measurements showed gain indicating that the beam was being refocused. Some rough estimates are in reasonable agreement with the observed bending of the plates. To minimize these effects, several measurements were made on each plate using both sides. The overall result indicated an average loss of about 5%, but with a standard deviation of about 5%. No systematic dependence on paint thickness was found. We conclude that there is probably some scattering due to the surface roughness of the paint, but it is only at the level of about 5%.

IV. Loss Measurement at 130-360 GHz

The National Physical Laboratory tested a separate set of samples using a Fourier Transform Interferometer. The results are included as an appendix. Although the frequency range is higher than will be used on the GBT, the measurements can be

February 28, 1992 Page 3

expected to place an upper limit on losses at lower frequencies. The average loss at 130 GHz is less than 3% and the maximum is 4.5%. The numbers are reasonably consistent with the 90 GHz measurements.

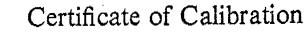
V. Conclusions

Dissipative loss in the paint samples is negligible at frequencies below 100 GHz or so. Some scattering is observed at the 5% level, but better measurements (e.g., on thicker plates which will remain flat) would need to be made to confirm this and to determine the angular region of scatter.

Phase shifts due to the dielectric should not be a problem even for 4 mil thicknesses up to ~150 GHz.

NATIONAL PHYSICAL LABORATORY

Teddington Middlesex TW11 0LW England





THE DETERMINATION OF THE POWER REFLECTION SPECTRA OF SOME PAINTED ALUMINIUM SHEETS

FOR:

National Radio Astronomy Observatory

Edgemont Road

Charlottesville

VA 22903, USA

CUSTOMER REFERENCE:

#POO131

DESCRIPTION:

6 painted sheets, approximately

 $100 \times 100 \times 1.7 \text{ mm}^3$

REFERENCE C15/0089

PAGE 1 OF 2

DATE

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16 December 1991

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NATIONAL PHYSICAL LABORATORY

Continuation of Certificate

MEASUREMENTS

The normal incidence power reflection spectra of six painted aluminium sheets were determined in the spectral region between 120 and 360 GHz by the technique of Fourier transform spectroscopy. The interferometer used in the measurements was a polarising wire grid instrument with a liquid helium cooled indium antimonide hot electron bolometer as its detector. The spectral resolution was approximately 30 GHz, and the measurement temperature 293 K. The geometry of the interferometer was such that the extreme off axis ray incident on the specimens would have had an angle of incidence of about 2°. The incident radiation was plane polarised.

The measurements were initially made using an alumined glass blank as the reference reflector. The reflection spectrum of this was subsequently determined in a separate comparison against the NPL standard of far infrared reflectivity based on a silicon disc. The calibrated spectrum of the reference reflector was then used to derive the reflection spectra of the specimens from the measurements.

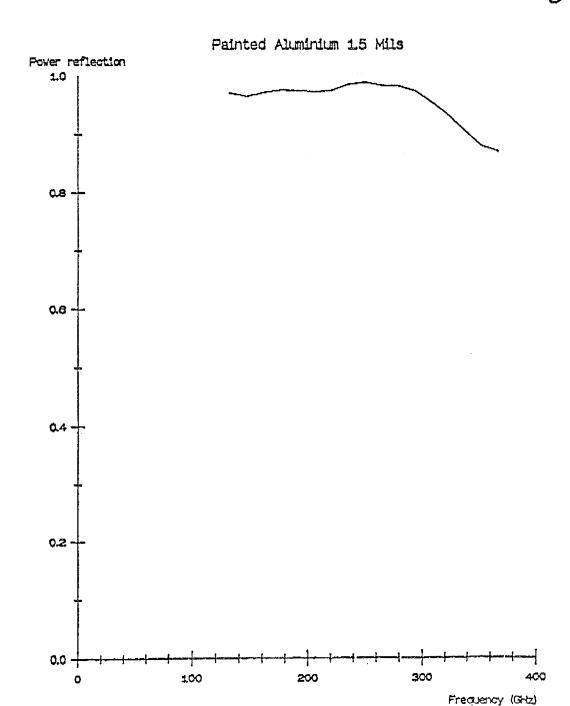
RESULTS

The results of the measurements are shown in figures 1 to 6. Each spectrum represents the average of four independent determinations. The reproducibility of these was such as to suggest a random uncertainty in each average reflection spectra of about 0.01, or less, over most of the measured spectral region.

In addition to the measurements described above, two of the specimens (3.0 and 4.0 mils) had their reflection spectra measured in 4 orientations, for orthogonal orientations of the incident polarisations on both painted surfaces. The results of this are presented in figures 7 and 8. These show significant differences, particularly for the 3.0 mils specimen.

The tabulated values of the spectra shown in figures 1 to 6 are presented in tables 1 to 6.

PAGE 2 OF 2



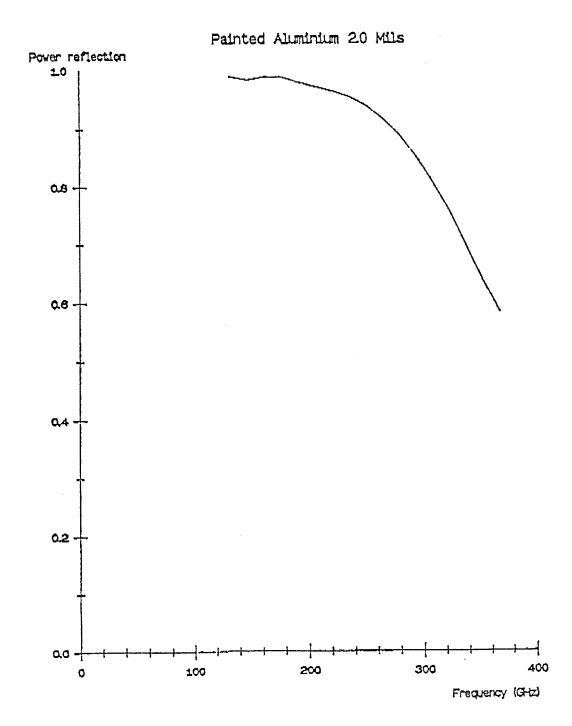
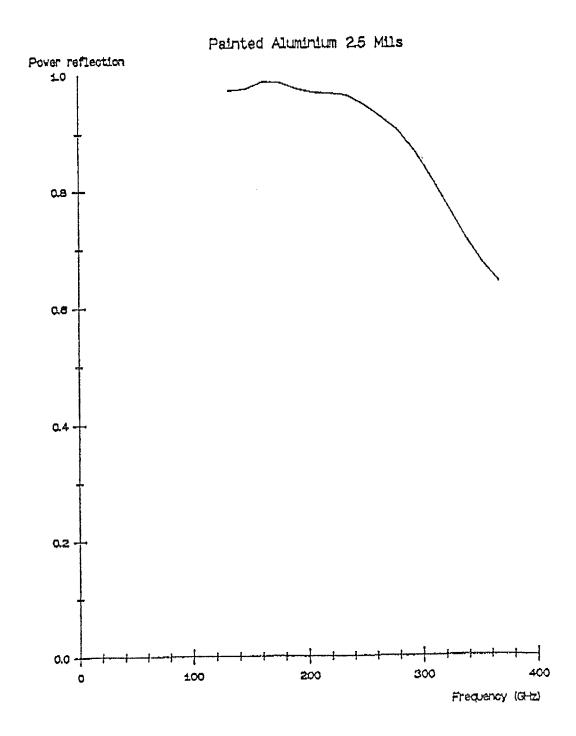
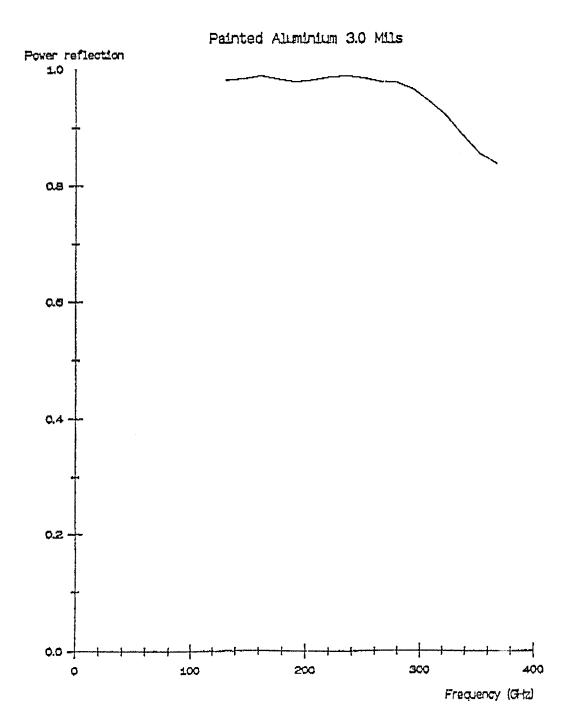
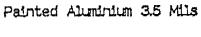
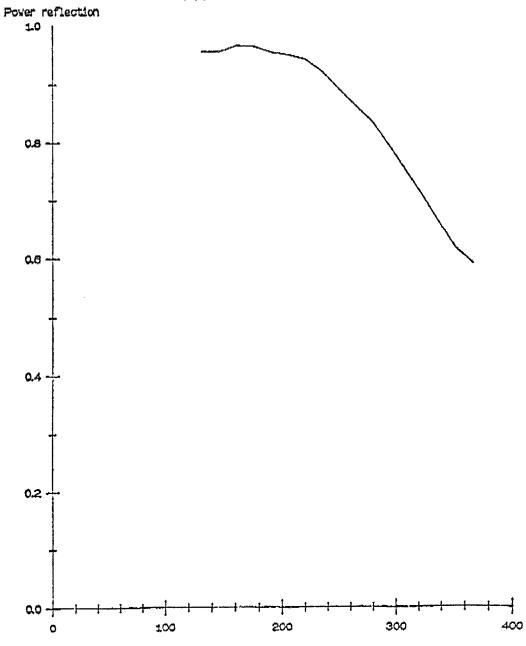


Fig 3.









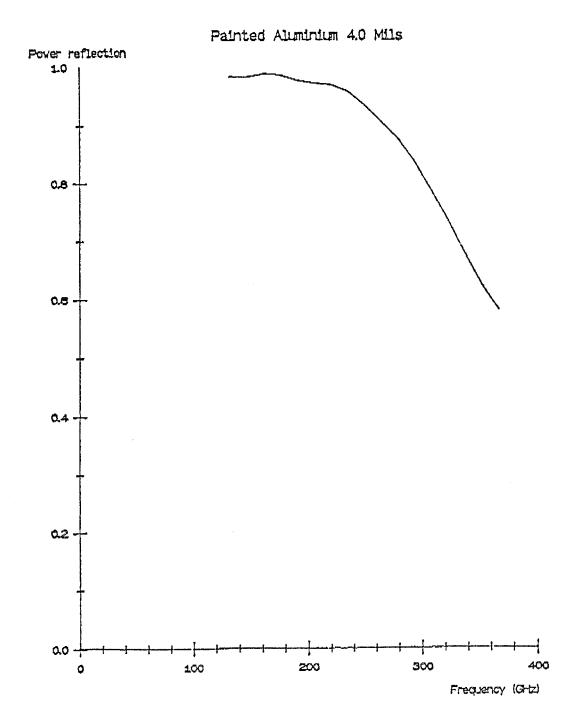
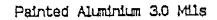


Fig 7.



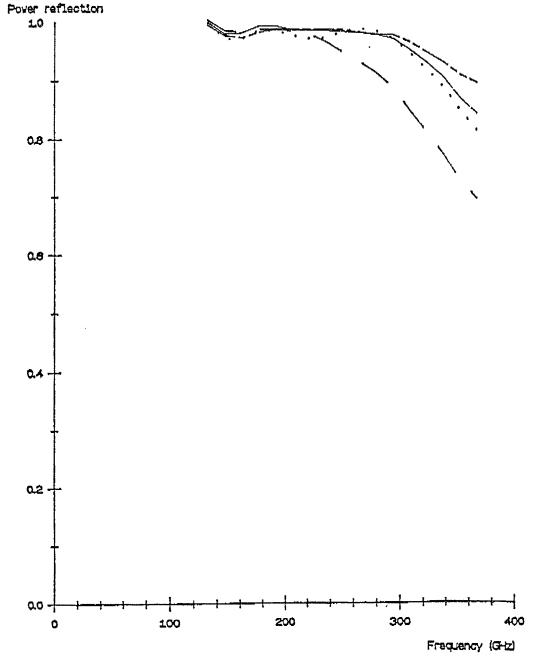


Fig 8. Same key as fig.7.

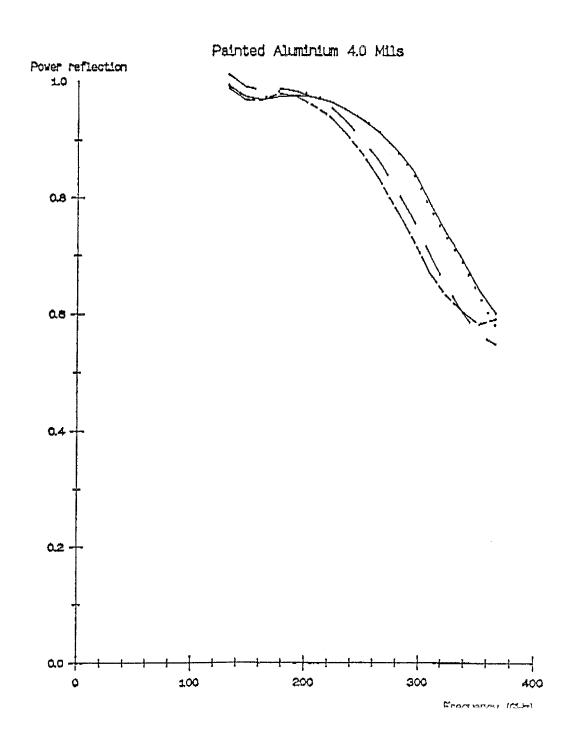


Table 1

Painted Aluminium 1.5 Mils

Frequency	power
(GHz)	reflection
131.8359	0.96890
146.4844	0.96355
161.1328	0.97051
175.7813	0.97466
190.4297	0.97321
205,0781	0.97112
219.7266	0.97351
234.3750	0.98418
249.0234	0.98790
263.6719	0.98254
278.3203	0.98105
292.9688	0.97309
307.6172	0.95299
322.2656	0.93101
336.9141	0.90408
351.5625	0.87825
366.2109	0.86813

Table 2

Painted Aluminium 2.0 Mils

Frequency (GHz)	power reflection
131.8359 146.4844 161.1328 175.7813 190.4297 205.0781 219.7266 234.3750 249.0234 263.6719 278.3203 292.9688 307.6172 322.2656 336.9141 351.5625	0.98761 0.98212 0.98773 0.98731 0.97850 0.97039 0.96294 0.95340 0.93896 0.91689 0.85082 0.85082 0.80553 0.75533 0.69598 0.63573
366.2109	0.58391

Table 3

Painted Aluminium 2.5 Mils

(GHz) reflecti
131.0359 0.9721 146.4844 0.9755 161.1328 0.9877 175.7813 0.9856 190.4297 0.9754 205.0781 0.9686
219.7266 0.9667 234.3750 0.9633 249.0234 0.9480 263.6719 0.9269 278.3203 0.9042 292.9688 0.8673
190.4297 0.975 ¹
146.4844 0.9755 161.1328 0.9877 175.7813 0.9856

Table 4

Painted Aluminium 3.0 Mils

power
reflection
0.98137
0.98336
0.98836
0.98248
0.97767
0.98084
0.98600
0.98823
0.98432
0.97828
0.97698
0.96524
0.94407
0.92034
0.88654
0.85416
0.83693

Table 5

Painted Aluminium 3.5 Mils

Frequency (GHz)	power reflection
	-
336.9141	0.66077
351.5625 366.2109	0.61634 0.58991
200.5103	0.50331

Table 6

Painted Aluminium 4.0 Mils

Frequency (GHz)	power reflection
	•
336.9141 351.5625	0.67690 0.62285
366.2109	0.57994