Lecture Notes

NRAO Lecture Program - Summer 1964

Review of

DATA ON PLANETS AND PLANETARY RADIO EMISSION

by

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	Mass (Mæ)	Mean Radius (R⊕)	Visual Albedo A	Т _М (°К)	Т _А (°К)	D _M (arc_sec)
Mercury	.0543	0.38	0.058	625	• • •	10
Venus	.8136	° 0.961	0.76	3 24	229	. 970
Earth	1.000	1.000	0.39	349	246	• •
Mars	0.1069	0.523	0.148	307	Ž17	10
Jupiter	318.35	10.97	0.51	145	102	50
Saturn	95.3	9.03	0.50	107	76	20
Uranus	14.54	3.72	0.66	69	49	4
Neptune	17.2	3.38	0.62	56	40	2
P1uto	0.033 (?)	0.45	0.16	60	42	0.1
••••						
Moon	0.0123	0.273	.072	387	274	1800

References (25) and (26)

Where D is the approximate maximum angular diameter of the planet as seen from earth, and

$$T_{M} = \begin{pmatrix} S_{\odot} & (1-A) \\ T_{M} & = \begin{pmatrix} T_{M} \\ T_{\odot} & \end{pmatrix} \qquad T_{A} = \frac{T_{M}}{\sqrt{2}}$$

where S_{a} = the solar constant

 r_{o} = mean distance of planet from the sum in A.U. σ = the Stefan-Boltzmann constant

 T_{M} is the equilibrium temperature of a perfectly insulating gray surface at a distance r from the sun, and is the highest temperature of a solid body of albedo A can attain as a result of solar radiation.

 T_a is the equilibrium temperature of a perfectly conducting gray sphere at a distance r from the sun, where the heat input is solar heating alone.

Most Recent Observed Infra-red Temperatures ($\lambda = 8-14 \mu$)

Mercury	613°K [†] (1)
Venus	215°K* (2) 235°K* (3)
Mars	Varies during Martian day between about
	205°K and 290°K (4)
Jupiter	128°K (2)
Saturn	93°K (5)

Moon $\sim 395^{\circ}$ K at "noon" to < 125° K at "midnight" (6)

(Numbers in parentheses refer to the reference number.)

* This temperature refers to the center of the disk. The average temperature over the disk would be $\sim 10^{\circ}$ lower for Venus and $\sim 3^{\circ}$ lower for Jupiter.

* Pettit did not have available modern interference filters for the isolation of the 8-14µ region. The filtering materials used, plus the atmosphere, transmit between about 3 and 14 μ . This temperature is the "noon" temperature on Mercury and refers to a phase angle of zero.

Radio Wavelength Data

Mercury

λ		T _B	Reference
3.45 and 3.75 (ĊŴ	~550 <u>+</u> 100°K	(7)

The above data refers to the range of phase angle between 60° and 110° , and it is not obvious from the data that the brightness temperature is a function of phase angle. However, the brightness temperature is much higher at this phase than the infra-red data gives, and might imply that at 3.5 cm the "noon" temperature reaches ~1100°K.

Venus

λ	T _B	Reference
.8 cm	315 + 70°K	(8)
.86	410 + 160	(9)
3.15	595 <u>+</u> 55	(10)
3.37	575 <u>+</u> 58	(11)
3.4	575 <u>+</u> 60	(12)
3.75	585 <u>+</u> 7	(13)
9.4	580 <u>+</u> 160	(10)
10.2	600 <u>+</u> 65	(14) and (12)
21	520 + 33	(15)
, 40	400 + 57	(15)

Besides the "roll-off" in the spectrum near 1.5 cm, Venus also exhibits a brightness temperature which is a function of phase angle, with the minimum brightness temperature occurring near inferior conjunction, but evidently not at inferior conjunction. The most accurate measure of this "phase effect" is probably that of Drake (15) at 10.0 cm, which gives

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 $T_{B} = 622 + 41 \cos [i + 21^{\circ}]^{\circ}K$

where the + (-) is used after (before) inferior conjunction. (This plusminus sign convention is due to the fact that the phase angle increases too, and reaches 180° at inferior conjunction, but decreases from 180° after inferior conjunction.)

Mars

Jupiter

,	λ		T B		Reference
	3 cm		218 <u>+</u> 50°K	nan fra de Branne, de trainformente de Carlos de C	(10)
· .	3.14		211 <u>+</u> 20°K		(16)
	10	. *	177 <u>+</u> 17°K		(15)

The observations indicate that Mars behaves like a perfectly ordinary, perhaps slightly rough, black-body.

λ	A B B	Reference
9.4 cm	$A + 1/2 B = 658 + 58^{\circ}K$	(17)
10.0	635 + 4 106 + 30	(18)

continued --

TB	Reference
<u>A</u> <u>B</u>	
A + 1/2 B = 800 + 100	(19)
2296 1465	(20)
A + 1/2 B = 3380 + 50	(21)
A + 1/2 B = 5500 + 740	(21)
-	2296 1465 A + $1/2$ B = $3380 + 50$

Jupiter has a distinctly non-thermal spectrum, as may be seen above, and, because of the fact that its brightness temperature is so high at long wavelengths, was the first planet detected by radio astronomers. In fact, over the radio region, the flux from Jupiter appears independently of wavelength, which gives a λ^2 increase of brightness temperature with wavelength. When Jupiter is observed with a linearly polarized antenna in the decimeter wavelength region, its brightness temperature is a function of rotational phase. This changing brightness temperature can be separated into an A component, which is independent of rotational phase, and a B component, which is the amplitude of the variable term.

In the decameter wavelength region, Jupiter has been observed to give bursts of radiation which, although they are most common near 18 Mc/s, have been observed between 4.8 Mc/s and 43 Mc/s. These bursts occur most frequently when certain longitudes are on the central meridian.

The origin of the decimeter radiation is evidently in synchrotron radiation from radiation belts surrounding Jupiter, which belts are like the Earth's Van Allen belts. The origin of the decameter radiation is not known.

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λ	$\mathbf{T}_{\mathbf{B}}$	Reference
3.45 cm	106 <u>+</u> 21°K	(22)
10.0	196 + 44	(23)
9.4	$\int 141 + 20 \stackrel{\rightarrow}{E}$ parallel to rot eq	q. (24)
	$213 + 18 \vec{E}$ perpendicular to	
	rot. eq.	

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Saturn

The earlier observations at 3.45 cm and 10.0 cm were interpreted as indicating that Saturn was a black-body radiator; however, the recent observations at 9.4 cm by the NRL group indicate the presence of non-thermal radiation, and show the rather surprising result that Saturn is brightest when the antenna E-vector is aligned with the rotational axis.

Neptune, Uranus, and Pluto have not been detected by radio astronomers -- which is not surprising -- because if they are black-body radiators they would give antenna temperatures of less than .002°K on the 300-foot antenna at 20 cm.

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