PROJECT 2.625
PULSAR SIGNAL PROCESSOR
MEMO NO. 1

### NATIONAL RADIO ASTRONOMY OBSERVATORY

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To:

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From:

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Subj:

Draft Specifications for a Proposed RFI-Cancelling Pulsar

Spectrometer

These draft specifications were discussed on 3 September 1981 by J. Taylor, M. Damashek, R. Fisher and W. Brundage. I hope I correctly cover the important topics.

### Time resolution:

0.1 to 2 millisec  $\Delta\tau$ .  $\mu$ s resolution is useful only at Arecibo.

### Total instantaneous bandwidth:

20 to 30 MHz is necessary.

Up to 100 MHz useful, but only if front-end does not overload or produce detectable intermodulation products.

### Frequency resolution:

10 to 100 kHz  $\Delta f$ .

#### Channel-to-channel isolation:

20 to 30 dB center to center of adjacent channel.

40 to 60 dB center to center of  $\pm 2$ ,  $\pm 3$ , ... channels.

# Dynamic range:

≥ 30 dB pre-detection.

≥ 10 dB post-detection.

#### Dual polarization:

Split into 2 banks for 2 polarizations.

Sum identical frequencies in 2 polarizations.

# De-dispersion:

0 to N ( $\Delta\tau$ ) both non-inverted and inverted frequency spectrum, where N ( $\Delta\tau$ ) = ? ...

### Dispersion searching:

Data tap for sum of 8  $\Delta f$  to 16  $\Delta f$  after impulse blanking and after spectral (channel) blanking.

# Impulse blanking:

Each channel blanks time-domain impulse RFI by selectable criteria which do not blank pulsar.

# Spectral blanking:

Each channel blanks "continuous" spectral feature by comparison to adjacent channels per selectable criteria which do not blank desired spectral lines.

# Amplitude calibration:

Pulsed noise calibration controlled by spectrometer.

### Input IF frequency range:

50-500 MHz.

#### Output:

Direct to Modcomp for direct tape writing and/or further processing.

### Possible implementation:

Filter bank;  $256 \times 100 \text{ kHz/2}$  (128 x 100 kHz).

Time-domain signal-micro-processor on each channel.

Mini-computer spectral/time-domain/control processor; equivalent/compatible to Modcomp II.

#### WDB/cjd