NATIONAL ASTRONOMY AND IONOSPHERE CENTER

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Telex: 932454

June 14, 1982

Dr. Richard J. Fisher N.R.A.O. P. O. Box 2 Green Bank, WV 24944

Dear Rick:

I am sorry I do not remember the names of the other persons in the telephone conference on FFT machines that we had some time ago, so I address the letter to you.

I contacted Dr. Noble Powell of General Electric, Syracuse, New York, on building a 64 point real or complex machine without any of the interfaces (which we would supply). Data rate of 8 bit words would be 20 MHz max., settable to lower rates only. Multiplying coefficients would be of 8 bit accuracy. His group built the 10 MHz 1 K FFT transformer that Ivan Linscott used in his controversial experiment. They have now developed a new custom integrated circuit that runs at up to 10-15 MHz. It does one butterfly operation. He parallels a lot of his operations so he is able to run faster. All the 200 IC's of the machine are on one large Unfortunately, his asking price (\$70 K to \$100 K) is too much for us, and I think it is not justified. Most of the cost is in design salaries: he would provide 2 machines for the same price as one (or nearly so). If you are interested we could explore sharing the cost, but I think the design from scratch seems much more attractive, particularly since extending the bandwidth beyond 20 MHz is rather difficult in the G.E. design.

I also promised you some values of delays vs. frequency for some high Dispersion Measure pulsars. Enclosed please find copies of printouts for the two extreme cases: PSR 0950+08 (DM = 3.0 pc cm⁻³) and PSR 1859+03 (DM = 402.0 pc cm⁻³), giving the sweep rate (S) (MHz/sec) at each True Center Frequency, and its inverse in Millisec/MHz (which is proportional to the Dispersion Measure). Also given is the delay between two 20 KHz points located at the True Center Frequency (also proportional to the Dispersion Measure). As we discussed over the telephone the most desirable single channel bandwidth for best S/N ratio is equal to:

$$B = S^{1/2} \times 10^{-3}$$

where B is in MHz and S is the sweep rate in MHz/sec. This comes from equating the transit time through one filter bandpass with

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the risetime of that filter. However, for strong pulsars one may use other bandwidths when studying special events, or other reasons. I hope this information helps you. We will start FFT design work in the late summer. We are interested in keeping a collaboration running with you, so let us know your plans and let us exchange ideas.

Thank you.

Yours,

Valentin Boriakoff

V. BorraRdl (CD)

vb/it

cc: Martha Haynes

Mike Balister

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