# Quarterly Status Update (QSU) July – September 2010





Atacama Large Millimeter/submillimeter Array Expanded Very Large Array Robert C. Byrd Green Bank Telescope Very Long Baseline Array





The format followed on this agenda provides orientation to the structure of this briefing, gives some high level science results and metrics, and then reviews Observatory Science Operations, Site Specific Activities, and then Observatory-wide operations.

### **GBT Science Results GBT** demonstrates new technique for studying dark energy

Using the technique of 'intensity mapping', the GBT has made the first detection of HI 21 cm emission at substantial redshift ( $z \sim 1$ ). This technique involves low spatial and spectral resolution imaging to detect the aggregate 21 cm signal from thousands of galaxies at high z. Cross correlation with optical redshift surveys reveals a robust detection of the mean HI signal from these galaxies. This technique has the promise of mapping large scale structure at high redshift, such as the Baryon Acoustic Oscillations, and hence for determining the nature of the Dark Energy that drives cosmic acceleration. (Chang ea. Nature, 2010, 466, 463)

NRAO



### **GBT Science Results GBT/Mustang** makes highest SZ image to date

The MUSTANG bolometer array operating at 3.3 mm on the GBT has been used to make the highest angular resolution map to date of the Sunyaev-Zel'dovich effect (SZE) in a cluster of galaxies. The SZE map of the cluster, RXJ1347-1145, made at 10" resolution, confirms the presence of a localized enhancement 20" from the center of the X-ray emission, a feature that is interpreted as an ongoing major merger event. The GBT data also detect a pronounced asymmetry in the projected cluster pressure profile that would not have been detectable in lower resolution observations. With the recent improvements in the GBT surface, high resolution SZE measurements will become a powerful tool for studying intracluster gas.



Caption: Contours of the SZE decrement from MUSTANG superposed on the Chandra X-ray count-rate image smoothed to 10" resolution. The anomalous SZE enhancement is the ridge to the left and below the central X-ray source.





# **EVLA Science Results** First EVLA publications II:

Extended, low excitation molecular gas in submm galaxies at  $z \sim 2$ 

The EVLA has revealed low excitation molecular gas in active star forming galaxies at  $z\sim2$ , during the 'epoch of galaxy assembly.' These observations show more total gas, by a factor 2, than previously derived from higher order transitions, and indicate that the gas is extended on scales  $\sim 16$  kpc.These results challenge many preconceptions on massive galaxy formation at high z. (Ivison ea 2010, MNRAS, in press).

NRAO



#### **EVLA Science Results** VLBA imaging of neutral atomic gas within Ipc of a massive black hole:

The VLBA has been used to image the atomic Hydrogen on sub-pc scales near the active nucleus of Centaurus A galaxy. The HI absorption lines are seen against the radio loud nucleus and jet. It is found that the broadest HI components are located within 1 pc of the nucleus, as opposed to being intervening clouds further out in the di The result indicates that dissociation of H<sub>2</sub> near the nucleis efficient, and may foster the fueling process of the supermassive black hole (Espada ea. 2010 ApJ 720, 666).

NRAO



Figure:VLBA image and HI 21 cm absorption spectra toward Cen A.



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**Scheduled** = planned observing time.

Astronomy = amount of observing hours that concluded Downtime = amount of hours lost during observing

**Maintenance** = scheduled period for technicians to service. Observing time is not scheduled during this time. This time is considered 'protected' and is not interrupted for targets of observing opportunity.

**Unscheduled** = time that went idle (unplanned); for example, for VLBA if no media was available or due to the 10 weather environments and the tiger team visits; for VLA if no dynamic project fit, for GBT = holiday.

**Downtime** = faults that occur during a planned observation; e.g., circuit breaker fault, fraction of array unavailable, etc.



All metrics are compiled by principal investigator, not project team.

Top graph is in **observing hours**. Bottom graph is in % **of observing hours**.



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This information is obtained from the proposal coversheet which includes scientific categories. The proposals tend to include one to three scientific categories per project. The metrics are created by splitting time (minutes) evenly over the categories listed on the proposal coversheet.

Basic analysis (some trending may be due to seasonal variations in activities; full year trending analysis will be performed at the culmination of the fiscal year):



No proposals solicited during this quarter.

During FY2011, proposals were on a four-month cycle (October 1, February 1, June 1). This will be changing to a six-month cycle in FY2011, beginning with Feb 1 2011.



I. Total data served from all web servers.

2&3. Specific plots for the main web site and for the science web site



Although the three month history does not demonstrate the change fully, this is a really impressive change from the past. Since the commissioning of the EVLA WIDAR correlator in March, the data served by the archive has burgeoned: increasing from 194GB in February to 16,074GB (that's 16.1 TeraBytes) in August, i.e., by more than 8,191% in 6 months!



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The resources (FTEs and costs) for OSO were included in the NRAO operating WBS for FY2011.

- There is a tool on the science web to **Coordinate and Consolidate meetings** for all subjects and telescopes.
- There has significant progress on the User Access & Support functions, including successful participation in the first and second rounds of ALMA integrated testing. Planning for the NRAO User Portal included a test installation of *Plone* and procedures for a mass migration of existing web content into a *Plone* environment.



Logging of metrics data is in place

NRAO Archive Strategy is now being executed, leveraging NGAS servers and the successful grant of 200TeraBytes of archive storage from TeraGrid for GBT Pulsar survey data.

Scalable storage systems on-line for EVLA, VLBA, ALMA and GBT data, although work remains on user access tools with focus on survey data for GBT. Backup archive system from Socorro to CV on-hold pending router upgrade in Nov 2011.

Resources at both NCSA (National Center for Supercomputing Applications @UIUC) and TACC (Texas Advanced Computing Center) have been committed for code development and data analysis with data made available through the TeraGrid network as needed.







The first graph illustrates the full lifecycle. The second graph is this fiscal year view. The vertical line represents where we are today. The CP represents the critical path.



MANAGEMENT: There are no activities to report in this period.

SITE: The delivery of **switchgears** continues and will be completed on schedule during October 2010. **The AOS Utilities Contract** was delayed compared to the schedule due to the contract termination on June 30, 2010. The expected date to restart of work is November 2010. A new bid process is proceeding. **AOS road construction** is ongoing, is 58% complete, and will restart in October after the winter break.

The **Electrical Infrastructure Implementation Review** conducted in Santiago 7-8 June 2010 to evaluate the overall status of all electrical infrastructure deliverables has action items that are being followed up.



ANTENNA: Additional structural modifications of the nutator were required as a result of the nutator analysis and testing done in Q3. Single pitch carbon fiber reinforced plastic (CFRP) was required to increase the overall stiffness of the subreflector and Invar counterweights were required to avoid surface deformation due to the mismatch in the coefficient of thermal expansion between the stainless steel counterweights and the CFRP mounts. This material had to be imported to Taiwan. All structural changes have been completed for units I & 2. The tuning of the nutator's servo control system continues to be a challenge. To speed completion of the tuning, one nutator system will be shipped to Green Bank, WV where the Electronics Division will work in parallel and collaboratively with the engineers at ASIAA who will retain nutator #2 to work with. The risk posed by the delayed delivery of the nutator is a delay in the ability of the Melco (EA) antennas to provide the total power measurements required by some science projects. The risk is being mitigated with a contingency plan that implements a fast scanning technique on the Melco antennas.

The **Production OPT (POPT)** Unit #1 was successfully used to perform acceptance testing on Vertex Antenna #10. The all-sky and offset pointing measurements were consistent with earlier testing with prototype OPT. Unit #1 will be transferred along with Antenna #10 to AIV for their use and Unit #2 will be used by AIPT for testing on Antenna #11 and on. A new objective lens design and construction was completed by colleagues at HIA using a lens mounting technique first developed for the GMOS instrument on the Gemini Telescope. This lens was tested and reduced the amount of temperature dependency by approximately 50%, but did not completely eliminate the drift. The remaining contributor has been found to be the CFRP optic tube. The existing tube is manufactured

using a 45° cross-hatch pattern of multiple layers of fabric. A new CFRP tube is going to be manufactured using unidirectional CFRP pre-pregnated material.

During Q4 FY2010, the 7th, 8th and 9th Vertex antennas were accepted into ALMA, and acceptance documentation for the 10<sup>th</sup> Vertex antenna is being finalized. Thus, the planned acceptance rate of 3 antennas per quarter was achieved. The acceptance of DV10 was delayed approximately 2 months due to the refurbishment of #1, the additional inspections of #1 to uncover root cause of surface degradation, and commissioning of Production OPT for use in acceptance testing. NA AIPT expects to deliver 2 antennas (10 & 11) in QI FY2011. A significant event in Q4 was the investigation and resolution of the root cause of the surface degradation first discovered on DV02 and later confirmed on DV01. The root cause was damaged panel adjusters that were overtorqued during installation. The over-torquing resulted in plastic deformation to the anchor nut of the adjuster and allowed the adjuster rod to move due to thermal expansion over many months. Vertex inspected 100% of the panels that had moved since the time the surface was initially set and found 21 of 50 adjusters were overstressed and had to be replaced. The correlation between degraded panels and damaged adjusters was 1:1. Lastly, a major effort to resolve a frequent communication problem between the ALMA computer and the Antenna controller was undertaken at the OSF by a team comprised of a Vertex software engineer and an ALMA computing engineer. The root cause was found to be in the CAN BUS driver used in the Vertex antenna. This driver software was rewritten and the problem has not reappeared.



FRONT END: The **NA FEIC leadership has transitioned** to John Effland and Mike Shannon, who will handle the technical and programmatic aspects, respectively. **FE Components, managed by S. Michalski, is now off the critical path.** 

**One integrated NA FE** underwent testing in Q4, but will be held in Charlottesville to be used for design verification tests. Three more front ends are in process of assembly and test in parallel with the design verification, with delivery of two scheduled for Q1 of FY2011. The delay in shipping will reduce the cost risk of any design modifications but may cause an overload of the AIV process as a large number of FEs will be delivered in rapid succession. **FE LO and test source production** (which compete for resources) are keeping up with the project's needs. **Technical problems with Band 7 WCAs were solved** and full production mode for all four primary bands in underway.



FRONT END: FY2011. The **Band 6 cold cartridge assembly (CCA) is in full production** with the formal approval of the specification change for the cross polarization. **Band 3 CCA** is having problems meeting its cross polarization specification. A **specification change for the cross polarization for Band 3 similar to the Band 6** was submitted. **The remaining FE purchase contracts** were placed and the difficulty of our suppliers not being able to obtain critical components has been overcome and delivery is proceeding as scheduled.



BACK END: Antenna Articles (AAs) have continued to be integrated in North America according to schedule with AAs 31-40 crated for shipment in August 2010. The next batch of 10 will be finished and ready for shipment in December 2010. These 20 AAs will be stored in the VLA warehouse until early 2011 in order to maintain warehouse space at the OSF and to minimize costs of any unanticipated modifications to the deliverables. They can be shipped at a moment's notice. The remaining 16 AAs are shipping in mid-2011. Line Length Corrector (LLC) and Sub Array Switch (SAS) LRUs, 16 each, were delivered and installed in June and July of 2010 at the AOS. This installation completes the outfitting of the Central LO Article 1, and brings it to its full capability of supporting reference timing for 16 antennas. Looking ahead, the first article units for the Photonic Reference Distribution LRUs arrive in late Q1 FY2011 and are undergoing acceptance testing. These items are of interest in that they represent the last new components integrated into the Central LO Article 2 (CLOA2) which is the final photonics deliverable scheduled for installation in Q2/Q3 FY2011. All other components in CLOA2 are add-ons to those in CLOA1 and represent a lower risk.



CORRELATOR: In order to meet the project need for **operating 2 quadrants of the Correlator** simultaneously by April 2010, the delivery plan for quadrant 3 was changed to an earlier date so that engineering tests using quadrants 2 and 3 can be used to verify the 2-quadrant operation while quadrant 1 is used for AIV/CSV activities. Quadrant 4 construction was completed so that a software test bed remains available in Charlottesville. Quadrant 3 was shipped to the AOS TB, reassembled, and passed PAS review. All correlator modes have now been verified.

COMPUTING: The major **problem in the NRAO R7.1.1 portion of the patch** was related to the Data Capture component which assembles the data before submission to the Archive. This component was completely reimplemented (the previous implementation was not maintainable in the long term) and more problems than anticipated were discovered in it after deployment. The major lesson learned however was that the testing needs to include critical CSV developed scripts, including data analysis of the observed data. These will be identified and tested for subsequent releases, including R8. (Globally the major issue this quarter with CIPT software is in the Archive subsystem, which is an European deliverable.)

SCIENCE: ALMA Commissioning started 22 January 2010. The highlights of Q4 included producing a **test image from six antennas at the Array Operations Site** in August 2010. ALMA imaging ability, which goes as the square of the number of antennas, grew through the period as the **AOS array increased to eight antennas**. The eighth antenna was brought into the array the evening of September 30<sup>th</sup>, with test data acquisition following soon after. The twenty eight baselines available with eight antennas offers a substantial advantage in ALMA's ability to accurately represent emission from the

sky - there were ten baselines available at the beginning of Q3. Project Scientist Richard Hills wrote: "We also have data on [NGC253 carbon monoxide emission] at Band 6 (230 GHz) and Band 7 (345 GHz). The band 7 data has also been processed to make a map and [t]here is a plot showing the velocity gradient. The band 6 data was taken with 6 antennas giving 15 baselines and the quality looks good, but at present we only have a short track and we need to take more data before we can make satisfactory test images."



JAPAN PARTNERSHIP (SPO-7) First pre-production Band 4 and Band 8 Cold Cartridge Assemblies (CCA) were integrated and tested in NA FE#3. It was determined that Band 4 will need magnets added to the SIS mixer blocks to suppress Josephson current and that Band 8 will need filtered connectors on the cold IF amplifiers. The change control board rejected waivers for these two defects, and these cartridges will be replaced as some point in the future. **Development of an LO driver/frequency multiplier combination for Band 10** with the new designs presently in fabrication. The prototype that was delivered last quarter failed and was repaired and returned. A new design power amplifier chip for Band 10 that was scheduled for fabrication in Q4 will not be ready for tests until Q2 of next FY. **Frequency multipliers and WCA LO drivers were delivered** to cold cartridge manufacturers for use with all bands. A **Band 4 and 8 LO CDR** was held, and designs were released for production. Components for assembling Front Ends, including Band 6 cold cartridges, were delivered to all integration centers. The NA FEIC assisted the other integration centers with assembly and test of Front Ends. Support was provided for integrating Front End assemblies into antennas. Components for assembling the **EA FEIC FE assemblies** were delivered.



Both graphs show the **NSF budget allocation**. In the case of the overall plan, the cumulative allocation is the allocation actually provided by NSF plus the planned allocations in FY10 to FY12. For the **FY10 graph**, the allocation is the planned FY10 allocation of 42.76 M\$. The actual costs shown are the inception to date expenses for the bilateral project, as booked in the general ledger. These costs do not include the commitments. The FY10 graph shows that expense remains below the total available allocation.



OFFICE OF CHILEAN AFFAIRS (OCA): One expatriate arrived (Lewis Ball, ALMA Deputy Director). OCA has reviewed and signed a total of 12 new ALMA Local Staff Member contracts in the quarter, bringing the total number of employees for which OCA provides ALMA with legal, payroll and travel support to 225 local staff. Of these staff, 25 are assigned exclusively for AUI/NRAO activities. It was agreed that AUI's local employment of local staff be reviewed after three years to evaluate its compliance with Chilean legal matters, the cost efficiency of its operations, and responsiveness to programmatic requirements. This review was conducted 1-3 September 2010 and concluded "... overall that AUI has met the legal requirements for employing staff in Chile under the law that governs international observatories; established an effective, professional HR organization in Chile; hired over 80 local staff members; complied with existing management agreements among AUI, ESO, NAOJ, and the JAO; and successfully negotiated a three-year collective bargaining agreement with the newly formed employees' union. The committee compliments AUI on these successes..." The committee identified a number of issues that remain to be resolved and other issues that ALMA will face as it continues to grow and make the transition from construction to operations. The committee presented 20 recommendations for follow-up action.

OCA was heavily involved in the **union negotiations** that ended successfully with the signature of the final offer for a 3-year period on 12 August 2010.

Although in May 2010 a new OCA office space was found, subsequent visits to it revealed flaws in the design and restrictions in building and electrical installations that lead to **cancelling this as a viable option**. The problems found were: (a) leaky roof with HVAC equipment affected by the earthquake;

this situation was fixed by the landlord, however while doing so other problems became apparent: (b) the total area subject to the rent agreement was not all constructible as it was common property and that in order to use it required performing a rental procedure that would be on the edge of legality with respect to Municipal law. And finally, (c) the other renters of the building made us aware that the electrical connection to the floor was underspecified for a floor of our size, which we confirmed, requiring an upgrade from the electrical company. A **new office building was found in the area** (293 square meters in area, ~700 meters away from the new ALMA Santiago Central Offices), and negotiations are under way. **Outfitting will occur during in QI/Q2 FY2011**.

OCA has provided the legal and institutional support for contracts and procurements for ALMA as follows: a total of **120 purchase orders were made for ALMA Construction (941 k\$)** and **194 for ALMA Operations (JAO) (869 k\$)**. The activities for ALMA Construction involve those described in the Site IPT section, namely AOS Roads Construction Contract, AOS Utilities – Electrical and FO cables installation contract, Fiber Optic Cable supply and Contractors' Camp expansion (JAO activity). Monthly reports were issued to CONAMA (environmental authority) related to flora/fauna and archaeological follow-ups. The **termination of the AOS Utilities – Electrical and FO cables installation contract** involved additional activities required to legally close the contract and follow-up.



Staffing: **Two of the three open user support scientist positions were filled**. Scott Schnee of the Herzberg Institute, Victoria, Canada, will fill a NAASC user support scientist position, October I 2010 (Assistant Scientist/Astronomer). Adam Leroy, currently a Hubble Fellow at NRAO, will fill a NAASC user support position, March I 2011 (Assistant Astronomer).

Hardware and Software: The first NA ALMA Mirror **archive racks and Oracle database** servers were installed and improvements to the archive room cooling system were completed. An instance of the **ALMA helpdesk was deployed** in preparation for integrated testing of the observation planning software, and it was configured for Single Sign-On. Integrated tests of the **observation planning software**, from Helpdesk questions through to technical assessments and mock proposal review meetings were performed together with the JAO and the other ARCs. **Pipeline Heuristics User Test 6 and tests of CASA updates** were completed.

Workshops and Meetings: A special session, "Preparing for ALMA", was held at the May AAS meeting and a session on Observing with ALMA was scheduled for the January 2011 AAS meeting in Seattle. Several data reduction and Simdata guides were developed for CASAguides in preparation for the NRAO Synthesis Imaging School. CASA and Simdata tutorials and Observing tool and Splatalogue walkthroughs were given at the NRAO Synthesis Imaging School, and an ALMA Townhall was held.






Status, Sept 30, 2010:

Slide reflects:

Milestones 17 has been satisfied

Milestones 10, 11, 13 and 16 have been slightly delayed.

The bottom graph illustrates the full lifecycle. The top graph is this fiscal year view. The vertical line represents where we are today. The CP represents the critical path.



Spending in FY10 is slightly behind schedule due to the delayed deployment of X-band receivers (\$205K) and 3-bit sampler modules (\$317K).

Financials are reported through September 2010. The FY10 graph is different from that in the Q1 report due to an error in fiscal reporting for Q1.

The FY10 budget/cost figure shows funds allocated for expenditure in FY10, only. Additional funds are on hand, but they are assigned to project activities, such as receiver production and installation, to be completed in FY11 and FY12 and to the retirement of project risk. The amount of these additional funds currently totals about \$2M.



The installation of the X-band horns exceeds the rate of receiver installations.

The time required to select the OMT design for the X-band receiver delayed its deployment start date to August 2010. However, the production and installation of the receivers are still scheduled for completion in December 2012.

Other EVLA receivers include L- (1-2 GHz), S- (2-4 GHz), and C-bands (4-8 GHz).

Impact of late Ka-band Rx should be slight since only 27 antennas/receivers can be used at a time. The  $27^{th}$  Rx to be installed in Dec;  $28^{th}$  Rx to be installed in Feb 2011.



The late deployment of the 3-bit, 4Gsps samplers has delayed the availability of 8GHz observing capability, but it does not delay the overall completion of the EVLA project.



WIDAR baseline boards:

Final four baseline boards were delivered in September, 2010.



EVLA Data Access: Dataset sizes obtained through the Open Shared Risk Observing program are up to an order of magnitude larger than any obtained previously be the VLA. These are currently being made available to the user community through an Archive Access Tool and ftp, via a fast I Gbps link from the Domenici Science Operations Center to the Internet-2 hub in Albuquerque. Further improvements in the way archive queries are handled have increased the speed of data access during Q4. While in theory the I Gbps link will have the bandwidth to support ftp downloads of future wideband data (in mid-2011 the data rates will increase by another order of magnitude), in practice the speed with which users can download data will probably be limited by general network traffic and the link speed at the recipient's end. We have therefore begun planning an alternative model for the dissemination of large datasets, involving the shipping of hard drives to users upon request. This model will be tested with Resident Shared Risk Observing data during FY 2011.

Shared Risk Observing: Access to the EVLA Early Science is provided by two shared risk observing programs for the user community: the Open Shared Risk Observing (OSRO) program, and the Resident Shared Risk Observing (RSRO) program, along with one for EVLA Commissioning Staff Observing (ECSO). OSRO projects have comprised the majority of observing time on the EVLA. RSRO observations with 16 correlator sub-bands began in July. The D-configuration was extended to accommodate wideband science. The move to the DnC hybrid configuration took place in September.



EVLA Commissioning and Science Verification: The commissioning effort in Q3 focused on (1) improving overall system stability, (2) testing and verifying the new RSRO wideband observing modes and science data, and (3) testing the performance of the VLA's 74 MHz dipoles with the wideband EVLA electronics in preparation for the September 15 Call for Proposals.

Proposal Submission: The Proposal Submission Tool and the EVLA Observational Status Summary were updated to support the September 15 Call for Proposals, which included a special call to use the VLA's 74 MHz system.

PASEO: In July NM Operations hosted the first meeting of the Panel to Advise on Science and EVLA Operations in Socorro. The recommendations of this committee will be taken under advisement and implemented as appropriate.



ARRA funds use: VSQ sprinkler replacement is to be replaced by retirement of VSQs as a result of mouse infestation. NSF approved alternate use of these funds in Q4.

Management, Facilities, Telescope Operations, Hardware, and Software are covered under the EVLA sections.

Computer Infrastructure: NRAO internal work has been completed for this milestone. However, New Mexico Tech was having trouble providing promised connectivity to Internet 2 at full I Gbps bandwidth. This problem was resolved in April 2010.

Engineering Services: Ongoing work items include 5000 ties being replaced along the ~44 miles of array tracks by the end of Q4. Only 3000 completed by Q4 due to additional work need to repair intersections and the wait for tie delivery (used ARRA funds and no ties ordered in previous FY). The planned azimuth bearing change on antenna #25 was performed on antenna #6 instead, to coordinate with the repair of its azimuth gear box. Antenna #25 will have its azimuth bearing changed at a later date.

The reconfigurations into C- and B-configurations have been moved to FY2011. RISK: none; no impact on EVLA schedule.





**DSS:** Work continues on interim enhancement releases and the next major capabilities release in early QIFYII (2-week delay). Software effort this quarter was devoted to the completion of our scheduling algorithm infrastructure expansion and addition of a sensitivity calculator.

#### PTCS: NOTE: Beginning Q2FY10, for clarity, PTCS is tracked on this page as two subprojects

<u>PTCS-Surface</u>: The surface RMS goals for the GBT main reflector in benign conditions was met in Q3FY10. All holography maps to date have been re-processed with an improved formula for the correlated amplitude. This is expected to yield a slightly improved surface, and verification tests with Mustang are planned in Q1FY11.

<u>PTCS-Servo:</u> Acceptance testing for the digital servo has been rescheduled for Q4FYII due to operational calls on project personnel. Risk: Advanced modeling for servo is delayed, additional operational funds will be required once the Lockheed/Martin funds are depleted to finish the project. Mitigation: The schedule to begin work on model-based servo control will be delayed.

**CICADA (GUPPI):** The planned initial release of the GUPPI de-dispersion modes has been completed. Release of additional modes and ease-of-use improvements will now be managed through routine Green Bank instrument enhancement protocols.

#### **Camera Development:**

<u>MUSTANG100:</u> For FY10 we have established the MUSTANG100 project to track the potential upgrade to a 100-pixel array in the existing MUSTANG receiver. The delivery of a new array by

NIST occurred in Q4FY10 as rescheduled but a wirebonding issue compromised the first test run.. Risk: MUSTANG100 will not be ready for the upcoming high frequency observing season. Mitigation: Use Mustang 64 for the early part of the season. Note: The schedule for the 100-pixel array is in NIST, not NRAO control (NIST is donating the array).

KFPA: Final commissioning prior to shared risk observing in QIFYII was successful.

<u>Note:</u> In the Program Operating Plan, the milestones #7 and #8 were reversed. Mustang 64 was released for general use in QIFY10; KFPA commissioning begins in Q2FY10



Overall development expenses for Green Bank tracked the budget very closely for FY10. The **DSS** project finished the year 8% over budget due to additional contract labor added to keep the project on schedule while key vacant positions were in recruitment. **PTCS** was 41% over budget due to 1) overlap of scientists as Todd completed the surface work, transitioning off to ALMA slower than projected, 2) addition of a new project scientist to the project and, 3) the unbudgeted extensive use of two technicians to built multiple prototype motor interface cards for the PTCS servo project. However, due to lower expenses than projected for the other projects, the Green Bank development expenses finished 1% under budget at the close of Q4FY10.



## DSS:

A large portion of the software effort this quarter was devoted to completing the scheduling algorithm infrastructure to include hooks for climate models, usage of historical weather data, and more realistic modeling of receiver temperatures. These expanded capabilities will be released early in QIFYII, immediately benefiting high-frequency observers. A sensitivity calculator, which will assist astronomers in determining observation durations based upon science objectives, was also developed and is expected to be completed next quarter. This tool will ultimately be available to astronomers as part of the proposal creation and submission process to streamline proposal preparation. Likewise, we expect this tool to make the technical review process simpler as well.

## PTCS:

<u>Surface:</u> All holography maps to date have been reprocessed with an improved formula for the correlated amplitude. This should yield a slightly improved surface, and tests with Mustang are pending in the next quarter. Work continued on the GBT memo on surface panel deformations, comparing manufacturer predictions with holography measurements under different environmental conditions. The group has begun assisting the Sardinia Radio Telescope (SRT) group in their initial attempts at commissioning their holography system. Using effort provided by the SRT and a U. Maryland graduate student (funded by AUI), we have started a new round of "out-of-focus" (OOF) holography data processing improvements.

<u>Pointing</u>: Measurements completed on July 1&2 and confirmed successful completion of the track project; published pointing corrections and required updating of pointing model are delayed due to change in PTCS staff. GB staff has developed a prototype system to use the Subreflector to correct for the effects

of pointing due to wind-induced deflections of the feedarm. This approach appears extremely promising, and will be converted into a production system as software resources become available.

<u>Servo:</u> Much of the activity on the PTCS servo system in the quarter surrounds the integration of the EtherCat control protocol into the servo lab's GBT motor simulator; which provides an environment to drive all of the simulated elevation and azimuth motors from the new digital servo control systems. All integration work for the servo subsystems will be tested on this environment before they are released for the GBT. Installation of servo hardware begins in Q1FY11 but due to staff members diverted to work on other important operational issues and unscheduled telescope maintenance the completion and acceptance testing schedule are delayed. The POP calls for the replacement of the current servo system with a digital system by the end of Q4, however this deployment is now delayed until Q4 FY2011. Risk: Advanced modeling for servo is delayed, additional operational funds will be required once Lockheed/Martin funds are depleted. Mitigation: None. Risk will be assumed. The performance at the various milestones is meeting spec and expectations while the schedule to achieve the associated milestones slips. The PTCS work is so fundamental to the GBT strategic future, it must be completed, even if operational funds are required.



## CICADA:

GUPPI: The planned initial release of the **GUPPI** de-dispersion modes has been completed. Release of additional modes and ongoing ease-of-use improvements will now be managed through routine Green Bank instrument enhancement protocols.

SPECTROMETER: **The GB spectrometer ATI grant** through UC-Berkeley has been funded. A project kickoff meeting was held on September 2, 2010 establishing a baseline set of milestones, action plan for the project, and initial specifications. NRAO and UC-Berkeley are in ongoing dialog to refine the specifications in accordance with the design phase as described in the ATI proposal.

NAIC PULSAR BACKEND: NRAO has agreed to build a version of its GUPPI **pulsar backend for the Arecibo telescope** in Puerto Rico. With a minimum amount of Green Bank effort, the Arecibo Observatory is able to leverage the significant work by NRAO and the Berkeley CASPER group for an affordable world-class backend to use in science related to timing microsecond pulsars and nanograv experiments.

#### CAMERA DEVELOPMENT:

MUSTANG100: **MUSTANG64** on the GBT and in use for regular observations, closing out that project. Per the POP, the array for this instrument will be replaced with a 100-pixel array produced by the National Institute of Standards and Technology (NIST). NIST completed a **100 pixel detector array**, using results from U.Penn's witness pixel measurements earlier in the summer, and installed it in the array package. A wirebonding issue, now being addressed, compromised the first test run. With the 90 GHz observing season imminent and a heavy proposal load, the original 64 pixel array was reinstalled and 100 pixel retesting will occur later in the season.

KFPA: The **KFPA receiver** received modifications to the cryogenics between commissioning runs reducing the amplifier physical temperature so the system temperatures for the receiver are comparable to the EVLA K band receivers and below the 35 K maximum specified system temperatures. The investigation into some spurs was complete and Config Tool requirements modified to reflect the LO settings required to mitigate any in-band spurs. All observing modes were tested and each of the seven beam calibrated. Integration into the GBT observing systems prepares the receiver for the shared risk observing call for QIFY11. Testing and documentation of the **KFPA Python data reduction pipeline** nears completion and position switched data reduction has been completely tested from observation planning, observation and automatic data reduction by a number of test observers.

W-Band FPA: Work progresses toward a conceptual design of the W-Band FPA receiver. Oversight of the R&D aspects of this work will be transferred in QIFYII out of the Green Bank Development auspices into the Coordinated Design Laboratory in Charlottesville.



A core team of scientists and engineers presented the design concepts for **the 4mm Two-pixel receiver** to a joint meeting of Green Bank and CDL management and staff. The conceptual designs were validated with the primary unresolved issue being the need for a cold isolator. To mitigate the risk of this unresolved element, the team has devised a plan to build a single pixel RF thread in the production dewar and perform baseline stability and polarization purity measurements to determine if the planned design requires the additional components.

New hybrids were installed in the **Ka correlation receiver** and the receiver was reinstalled on the GBT in Q4FY10 for Zpectrometer, Spectrometer, and Continuum backend characterizations. Initial results show baseline improvements since the rehabilitation. Once final characterization with the various backends is complete, Ka observers will be notified of the new performance limits and be invited to reinstate their projects.

Work has begun on rehabilitation of the **20m telescope** in Green Bank and two receivers; L-Band and X-Band. The frequency coverage of the L-Band receiver will be expanded to match the science requirements of the SkyNet project and NRAO will place the X-Band components in a new dewar for the better long-term cryogenic performance required of the **SkyNet** project.

Work continues on the **GBT VLBI Upgrade** with as much of the first round coding that could be started without the recorder complete and ready for testing. The Roach Digital Bank End is being returned to Socorro for reprogramming and will require some interface adjustments when it is returned. Two Green Bank staff will travel to Socorro in early QIFYII for implementation meetings

Work continues on data reduction and mining software to streamline publication of GBT data and science results.



In accordance with the rules for expenditures of stimulus funds, all parts and materials were purchased in Q4FY10 for all of the projects funded by the American Recovery and Reinvestment Act (**ARRA**). Four projects are complete and materials purchased for all others.

With additional painters contracted and generally cooperative weather this quarter, a greater portion of the **GBT was painted** than in earlier years. This is part of a multi-year increase in painting effort to bring the GBT into compliance with outside engineering consultant's recommendations for maintenance. Several surface panels were painted and then re-measured to verify that they remained within surface thickness specifications.

Frontier Communications has proposed to the State of WV and NRAO a path for a **broadband connection to WVU** which will ultimately connect to Internet 2 and/or LambdaRail. Technical capabilities will be reviewed with the stakeholders in QIFYII. Direct contact has also been made by NRAO with a second provider to evaluate alternative service options as a competitive and risk-reducing strategy.





Sensitivity Upgrade: Work on the fiber links to the PT and MK VLBA antennas was delayed by the delayed signing of the NRAO-USNO Memorandum of Understanding, which is providing the funding for this upgrade. The MoU was signed in Q4.

The trial installation of DBEs at Pie Town and Los Alamos were not met in Q2 due to collaborator delays (Haystack). The order of DBE/MK5 installations was modified to accommodate emergency maintenance work on the Pie Town antenna. By the end of FY 2010 four antennas were outfitted with the new recorder systems and recording at 2 Gbps has been demonstrated.



33 GHz receivers will be designed, built, and installed on the VLBA using ARRA funds if the MoU with NSF and USNO is signed.

8TB modules purchased have been unusable due to a firmware bug. Conduant Corportation will address this. Estimated entry into service of these 8 TB modules is Oct 20, 2010.



Draft MoU completed and awaiting signature by NASA and USNO. Both agencies committed to the development of the MoU, but found that they did not actually have committed funding when it was time to sign.

Azimuth wheel replacement at Hancock required because old wheel broke January 3, 2010. Major maintenance on HN, originally scheduled for Q3, was performed at this same time to avoid an additional trip.





Amplifier Production, Repair, and Development: Three prototypes of the P-band 230-470 MHz amplifier have been evaluated and tested. The mechanical design of the amplifier body of two production versions of the amplifier (they differ only by the hybrid used) is under way. Experimental evaluation of ALMA band #1 and #2 amplifiers awaits the availability of technician time. New amplifier production milestones included four 1-2 GHz, four 2-4 GHz amplifiers, four 8-18 GHz, and four 26-40 GHz amplifiers. Repair, upgrade, and retesting of amplifiers included one 1-2 GHz, two 8-18 GHz, four 18-26 GHz, one 26-40 GHz and four 38-50 GHz amplifiers. In total, 28 amplifiers were shipped. The EVLA amplifier production is slightly ahead of schedule. The deliveries of 18-26 GHz and 38-50 GHz amplifiers in support of MPI Receiver Group are on schedule.

*Electromagnetic support*: The Ku-band (12-18 GHz) phase shifters are scaled from Ka-band design; the X-band (8-12 GHz) is scaled from the W-band design. A total of 19 Ku-band and 2 X-band phase shifters were measured.



Amplifier Development: Research on general noise properties of three terminal active devices and in particular on noise properties of heterostructure bipolar transistors (HBTs), and CMOS MOSFET continues. A paper covering initial results has been published in IEEE Microwave Magazine (M.W. Pospieszalski, "Interpreting Transistor Noise," IEEE Microwave Magazine, vol. 11, no.6 pp. 61-69m Oct. 2010).

Using the THz SIS mixer test dewar, noise temperature improvement of the W-band 35nm MMIC LNA at lower ambient temperature was measured. Improvement of 4-6K in noise temperature was seen going from 17.5K to 8.5K ambient temperature, in agreement with theory. Bryerton gave a talk at Bonn *Receivers & Array Workshop* entitled "Cryogenic Performance of NGC 35nm InP Low-Noise Amplifier."

*Electromagnetic Support:* The S/X dichroic panel uses a <sup>1</sup>/<sub>4</sub>" Kevlar Honeycomb backing and lately this material is not available in small quantities. Aramid fiber honeycomb from Plascore Inc. was chosen as an alternate material after measurement and evaluation. One of two different designs of X-band phase shifter was chosen after measurements.



*Millimeter & Submillimeter-Wave Receiver Development:* Initial set of 700um SIS circuits had too thin AIN barriers, and therefore too high current densities. We therefore have stepped back to optimize the processing parameters, specifically the nitridation growth time, in order to accurately produce junctions with the correct current density. Three test wafers with different nitridation times are in the final steps of fabrication. Progress has also been made in direct sputtering of the AIN barrier (as opposed to using nitridation of AI layer). This is expected to give more repeatable current densities and is a necessary step in producing the NbTiN/AIN/NbTiN junctions, needed for quantum-limited SIS mixers up to 1.4 THz. The loss of the 400-500 GHz superconducting Nb hybrid did not go away upon cooling as expected. The resistivity of the substrate is the suspected culprit. The hybrid is being refabricated on high-resistivity silicon. In the continuing design of a 790-950 GHz SIS mixer, a Matlab program was developed to calculate the surface impedance of a normal metal in the extreme anomalous limits and calculated surface impedance values for 787-950 GHz for Nb, NbTiN and AI.

Advanced Receiver Development: The first **cryogenic Digital Orthomode Transducer (DOMT)** is being designed for S-Band (1.7-2.6 GHz) and will include a common-mode coaxial input for field-calibration. It is hoped that eliminating the cal coupler and injecting the calibration signal instead into an orthogonal mode, as well as minimizing the path length from the feed to the cryogenic amplifiers, will lead to a small but nonetheless beneficial reduction in noise temperature over our best current systems.

A **provisional patent application**, "Statistical Word Boundary Detection in Serialized Data Streams," has been filed concerning the techniques we have developed to implement low-overhead digital photonic links between receiver arrays in the field and backend signal processing facilities. The first

hardware demonstration of this concept will be done using low-cost PCB's, for which the schematics are nearly complete and the layouts are in progress.

The **35nm MMIC LNAs** are in fabrication at Northrop Grumman. We expect delivery of chips for testing early next year (no bullet on slide).



Phased Array Feed: Our BYU colleagues took our uncooled 19-element prototype array to Arecibo for tests on that telescope at Gregorian focus. These measurements will be of value in assessing the modified dipole for better noise match in the presence of mutual coupling. Our next measurement campaign on the 20-meter will include tests of both the uncooled array dipoles with better noise match and our first array with cooled LNAs. NRAO's new 40-channel data acquisition system with up to 5 Msa/s data streaming to disk will be used in these measurements.



The Precision Array to Probe the Epoch of Reionization (PAPER): Calibration of the 32 element Green Bank and South African arrays was completed. Data analysis is underway.

The engineering experiments such as the antenna rotation, receiver modeling, and study of ionospheric effects are graduate student projects.

Broadband Active Feed: The ambient temperature 300-3000 MHz feed is completed. Work continues on refinements of the cryogenic version.

LUNAR: The Lunar University Node for Astrophysics Research (LUNAR) is a grant from the NASA Lunar Science Institute to develop instrumentation for lunar-based research. Our current activity is centered around the Explorer-class DARE mission proposal with specific attention given to the front-end design concept.





NSPO 3.05 FTE in 2010 Q3 : Kellermann (1.0), Myers (0.25), Shepherd (0.9 with 0.5 paid by South Africa), Norrod (0.5), Webber (0.4 though has been spending most time on CDL/ALMA),

We have in principle agreed to manage the DVA-I as a 3-part project: (1) producing a design for a SKA Prototype Antenna as a deliverable of the TDP to SPDO, culminating in a CoDR around or before March I, 2011; (2) a costed preliminary design and plan for construction and testing of a DVA-I antenna likely at the EVLA site; (3) assuming resources are identified, actual construction and testing of DVA-I. There are milestones between each step, with opportunity for re-assessment. The DVA-I project management plan is being developed under the leadership of Chris Langley, who is targeted to take the role of DVA project manager assuming successful completion of the CoDR (and any other necessary design reviews).

The DVA-I Conceptual Design Review (CoDR) is now scheduled for Jan 2011 in Socorro, NM. The outcome of this first review of TDP deliverables will determine whether we move forward to prototype construction under TDP funding.





## **Social Networking**

• NRAO's Facebook fan numbers increased by over 27% during this quarter, from 1800 to 2300.

• NRAO's tweets have been "re-tweeted" (by recipients) 97 times during this quarter, reaching roughly 33,000 people PLUS the many thousands of people following the #astronomy, #seti, #space, and #womeinscience hashtags when they were appropriately used. One tweet that "went viral" with re-tweets entailed our pointing to a nice ALMA EPO resource posted on the ESO website; this also demonstrates NRAO's commitment to fostering success of the international ALMA EPO effort. (No thanks received from ESO after we told them we did it...)

## **News Media Activities**

• GBT intensity-mapping press release issued. Also directed several reporters to Chris Carilli for "independent commentary"

• Coordination with Chandra on Chandra/VLA release re 4C +00.58. (Changed spin axis of galaxy's SMBH)

- Coordination with Chandra on M87 "volcano" release
- SETI@50 press releases issued for GB public event and for Webcast of scientific workshop

- Interviews & coordination for Iowa Public Radio story on VLBA
- Prepared VLBA "talking points" memo for Walter Brisken radio interview

• Provided photos and information for upcoming (November) Popular Science item on ALMA, then did extensive interview with their fact-checker revising item

• Promoted VLA Open house and NRAO participation in M Mountain Fly-In during media interviews and at Socorro City Council meetings.

## **ALMA Explorer**

•The first in a planned series of "virtual tours" of NRAO facilities, the ALMA Explorer (http://www.nrao.edu/explorer/alma/) uses dozens of short hosted or narrated video clips, aligned atop high-resolution satellite imagery of the ALMA site, to introduce visitors to the facilities of ALMA, and its geographical and cultural contexts, all in a relaxed style akin to that of park way-finding signage or ranger tours.

#### **SETI** Workshop and Webcast

An educational, multidisciplinary workshop called "From Project Ozma to the Starship Enterprise, a Conversation about the Next 50 Years of SETI" (<u>http://www.gb.nrao.edu/ozma50/</u>) was held in Green Bank in September, in commemoration of the 50<sup>th</sup> anniversary of Project Ozma, the first-ever deliberate search for extraterrestrial signals of intelligent origin, which Frank Drake conducted in Green Bank using the 85-foot telescope. Fifty-three invited participants, including scientists, scholars, and communicators attended, including Frank Drake, SETI Pioneer; Dava Sobel, author of *Longitude* and *Galileo's Daughter*; Jon Lomberg, artist for the *Cosmos* Television Series with Carl Sagan; and Brother Guy Consolmagno, Meteorite Curator for the Vatican.

NRAO EPO staff broke new ground during the conference by increasing public participation in novel ways. Using a portable broadcast studio and the internet, NRAO EPO staff delivered a high quality, realtime, multicamera webcast of the conference via the free Ustream distribution service (<u>http://www.ustream.tv/</u>), which allows much higher numbers of simultaneous viewers than possible via an NRAO server. Through simple advertising on the NRAO Facebook page and Twitter feed we drew between 150 and 210 simultaneous viewers, who conducted their own parallel conversation (including NRAO staff interaction) using the Ustream chat feature, Twitter and Facebook. A blog was also developed to allow workshop participants to comment on potential topics of discussion. Now released to the public, the blog has received 1700 hits. See <u>http://ozmablog.gb.nrao.edu/</u>

A month long public celebration of the 50<sup>th</sup> anniversary of Project Ozma involved regional schools, county parks and businesses. Highlights included a public talk given by Dr. Frank Drake on September 11 which was attended by a capacity crowd of 150 people. In addition, the GB Science Center distributed Drake Equation Passports to county visitors who collected stamps by locating geocaches in our local state parks, and by visiting local businesses. Those who completed their passports won small prizes and were entered into a drawing for a telescope. About 100 people logged in to the official geocaching website to report their discoveries. An arts and letters contest for elementary/middle school children drew 90 entries, which were exibited at the GB Science Center.

## Filming Activity at the VLA

• Coordinated filming of NRAO employees for NM State University online feature about diversity in scientific/engineering workplace

• Coordinated filming at VLA for BBC documentary on "Do We Really Need the Moon?"

#### Diversity/Informal Education

• NM Highlands University is a certified Hispanic-Serving Institution.

#### **New Staff**

Sarah Scoles joined NRAO in Green Bank as our new education specialist, reporting to the Senior Education Officer. She holds a BA in Astrophysics and an MFA in creative writing and has re-located to Green Bank from her previous position at Cornell University.

## \$3K Grant

Received from the NM State Division of Tourism

#### Community event in Socorro

• Had solar telescope and other activities at a display at the M Mountain Fly-In and Socorro Aviation Day, an event that drew more than 800 visitors to Socorro Municipal Airport

#### Workshop on Science Writing

• Presented "case study" workshop with students of NMT Communicating in the Sciences course, on conveying scientific results to public. Course run by English professor for Physics grad students. Used "Hole in the Universe" paper and press release as case study.

Overnight Educational Events in Green Bank conducting research with the 40-foot telescope

•National Youth Science Camp (delegates from 50 states plus Central and South American delegates)

•Gross Pointe North High School Radio Astronomy Team (MI)

•AMRAD with NSF Spectrum Management Officer Andy Clegg

•Mt Vista Governor's School (VA)

•NOVAC Almost Heaven Star Party- Tour and overnight to use the 40 Foot (DC area)

•Civil Air Patrol Cadets (VA)

•Albemarle County Math, Engineering and Science Academy (VA)

•James Madison University (VA)

•Summersville Middle School (WV)

•Hampden-Sydney College (VA)

•Roanoke Catholic School (VA)

## Longer Events in Green Bank

•Society of Amateur Radio Astronomers Annual Conference

•Green Bank Star Quest Annual Multi-day Star Party with keynote speaker Carolyn Shoemaker. •Pulsar Search Collaboratory Teacher and Students Leader Institutes. 11 teachers from KY, MD, VA,

OH, PA, WV, MN, WI

•Undergraduate Students ( minority students through the LSAMP program) with the Center for Chemistry of the Universe

•WV Governor's School for Math and Science (60 rising high school freshmen from WV)

## **VLA Explorer**

•This will be the second in a series of web-based "virtual tours" of NRAO facilities. The program itself will be produced during FY2011, but filming was done in FY2010 Q4 in order to capture the VLA in its most photogenic D-configuration. Hosted by VLA scientist Rick Perley, the VLA Explorer will take

visitors into places that are strictly off-limits to the public (and to most of the staff!), including the WIDAR Correlator room and a VLA dish surface.

## NAASC Proposal Review

•Presented case for EPO component of NAASC proposal

# Astronomical Society of the Pacific (ASP) Conference

The ASP has become the meeting place for astronomy EPO professionals. This year's meeting was very useful as numerous case studies were presented on the effective use of social networking tools in EPO.




**Internal Communications:** A comprehensive internal communications plan was created and briefed to the AD team 30 Sep. COM organized the all-hands meetings and supporting multimedia content for the 29 July budget status meetings around the Observatory; and the NRAO functional alignment all-hands meetings that took across the NRAO on 5 October.

**Intranet redesign:** Work continued on a new NRAO Intranet site design that will transform the site into an effective internal communication tool. This work has been delayed somewhat by the resignation of Web Designer Taylor Johnson on 29 July. While we search for an individual to replace Johnson in this important full-time position, Johnson continues to support our internal and science community web design projects via a part-time consulting agreement. The internal web site design concept and structure was briefed to the AD team 30 Sep.

American Astronomical Society (AAS) meeting: In addition to the NRAO exhibition, the special events being supported by COM for the major winter 2011 AAS meeting (3000 attendees expected) include an NRAO Town Hall (Tuesday, 11 January, 6:30-8:30 pm), an ALMA Special Session (Wednesday, 12 January, 2-3:30 pm) titled "Observing with ALMA," an ALMA software tutorial splinter meeting (Wednesday evening, 12 January, time TBD), and an EVLA Special Session (Wednesday, 12 January, 10-11:30 am) titled "Early Science with the EVLA." Support materials being prepared for this AAS meeting include the 2011 NRAO Calendar, the 2011 Research Facilities brochure, and pre-loaded, branded 2 GB flash drives that will be distributed to NRAO Town Hall attendees.

American Association for the Advancement of Science (AAAS) symposium: A 90-minute VLBA-themed science symposium proposal titled "The Universe Revealed by High Resolution, High Precision Astronomy" was submitted in April 2010 for the 2011 AAAS Annual Meeting in Washington, D.C. This proposal has now been peer-reviewed, accepted, and scheduled for the 2011 AAAS Annual Meeting on Saturday, 19 February 2011 from 1:00-2:30 pm EST in the Walter E. Washington Convention Center, hall 146C. Our symposium's three speakers will be Mark Reid (Harvard Smithsonian CfA), Geoffrey Bower (UC-Berkeley), and James Braatz (NRAO). Mark Adams will organize and chair the session. NRAO Press Officer Dave Finley will coordinate media coverage with the AAAS Media Office.

**SC10 exhibition:** COM is collaborating with the CIS team to organize and lead the NRAO exhibition at the International Conference for High Performance Computing, Networking, and Storage & Analysis (SC10), which will be held at the New Orleans Convention Center, 15-18 November 2010. Approximately 11,000 scientists, engineers, software developers, CIOs, and IT administrators from universities, industry, and government agencies will attend.

**Science web site:** Collaborating with CIS on Plone content management system implementation. Q4 focus was establishing processes for bulk import of current science web site into Plone.

**NRAO Annual Report:** The initial design and text has been completed and reviewed with the Director's Office. The target audience includes funding agencies (NSF-AST et al), AUI, the AUI Board, Observatory Libraries and Directors around the world. A copy will also be provided to every NRAO/AUI employee. The report has been delayed by competing priorities, but is expected to go to the printer in November and be published in December. The Report covers FY 2010.





Green Bank

QZ administration database is being upgraded by NRAO IT personnel

AT&T cell phone service is available at Snowshoe and south of the GBT

Andy found that more than half of GBT proposals observe in a protected radio astronomy band New Mexico

Whitespace devices = unlicensed fixed or personal/portable devices that use locally-empty TV channels for such things as wireless broadband internet access. The first version of these rules excluded only a 2.4 km radius circle about the center of the wye



- Continental US
  - Hughes belatedly responded, favorably, to NRAO's letter of 2009 October noting FCC requirement to coordinate with NRAO regarding
  - installation of home earth stations within the National Radio Quiet Zone. Hughes and other installers have never instituted company-wide rules regarding compliance.
- International
  - 2010 meetings in Geneva finalized draft WRC input text regarding footnote protection of frequencies above 275 GHz
  - The IUCAF vice-chair authored a White Paper regarding preferred WRC actions for items of interest to radio astronomy
  - ITU-R RA. 1417 notes the importance of preserving the radio quiet environment of the L2 Sun-Earth Lagrange point





**Management activities** During the Jun 2010 call for proposals last quarter, we received a total of 235 proposals. The **GBT and the VLA/VLBA Proposal Selection Committees (PSC)** met in the second week of August to review the referee rankings and technical reports for all proposals with the goal of providing a recommended science program to the NRAO Director's Office. A list of the approved EVLA, GBT and VLBA/High Sensitivity Array (HSA) observing programs for Trimester 2010-C is now online (see <u>http://science.nrao.edu/observing/programs2010c.shtml</u>).

**Meetings** that were funded either partially or in full this quarter by the OSAA: From Project Ozma to the Starship Enterprise: A Conversation About the Next 50 Years of SETI, a workshop held in Green Bank from Sep 13-15 (see <a href="http://www.gb.nrao.edu/ozma50.shtml">http://www.gb.nrao.edu/ozma50.shtml</a>)

**NAASC** scientist hires were Adam Leroy as Assistant Astronomer, Tenure Track; Scott Schnee hired at Assistant Scientist/A, and now, Stuarrt Corder; **NAASC** postdocs are Amy Kimball, U. Washington; Robin Pulliam, U. Arizona; Nuria Marcelino, Lab for Molecular Astrophysics, Madrid.

**External funds:** \$10k from NSF to Karen O'Neil for the Project Ozma workshop; \$176.8k to Scot Ransom from the PIRE grant; \$43k to Rich Bradley from Cornell

NRAO received the following funds during this quarter for **grants awarded** to these members of the NRAO scientific staff:

-\$38,000, from NASA, Recipient Dale Frail, for "NRA/ Research Opportunities in Space and Earth Sciences" from Apr 2010 to Apr 2011.

-\$135,215, from NSF, Recipient Tim Bastian, for "Imaging Spectroscopy of Coherent Radio Bursts on the Sun: a New Probe of Magnetic Energy Release" from Jun 2010 to May 2013. -\$507,258, from NSF, Recipient *R. Craig Walker,* for "MRI-R2 VLBA Sensitivity" from Apr 2010 to Mar

2011.



**Summer Student Program:** For more information on the program go to http://science.nrao.edu/opportunities/summerstudents.shtml.

**Pre-Docs:** Cheng-Yu Kuo (Univ of Virginia Astr) continued working with Dr. Jim Braatz (NRAO/CV) on reducing and analyzing VLBI observations of water maser emission from galactic nuclei as part of the Megamaser Cosmology Project; and Josh Marvil (NMT) continued his appointment as a PreDoc this quarter working with Dr. Frazer Owen (NRAO/SO)

**Other Graduate Students:** Bin Chen (Univ Virginia Astr) continue to work under the supervision of Dr. Tim Bastian (NRAO/CV), funded through an NSF/AGS grant; Paul Ries (Univ Virginia Astr) continued to work under the supervision Dr. Todd Hunter (NRAO/CV), funded through the NRAO Student Observing Support program; Nicole Gugliucci (Univ Virginia Astr) continues to work under the supervision of Dr. Rich Bradley (NRAO/CDL), funded through the PAPER project; Rohit Gawande and Chatili Parashare (Univ Virginia EE) continue to work under the supervision of Dr. Rich Bradley



**Graduate Interns**: Charles Romero (UVA) is working with Brian Mason on quantifying the observational signatures uniquely accessible through high resolution SZE data (e.g., shocks and cold fronts, helium sedimentation), and applying them to MUSTANG data and if available data from GISMO on the IRAM 30m at 150 GHz; *Timothy Pennucci (UVA)* is working with Scott Ransom on the NANOGRav project; Sergio Dzib (UNAM) is working with Amy Mioduszewski on research projects related to the distance of young stars which will be part of his Ph.D. thesis, and *Leon Harding (NUI Galway)* is working with Greg Hallinan on broadband periodic dynamic spectra of ultracool dwarf pulsars.

**Undergraduate Interns**: Dana Sills, Cameron Welch, Aaron Cunningham, Deepak Rai, Sara Waters, James Durand, Jordan Leak, Scott Davidson (funded by EVLA), Matt Tibbetts (funded by EVLA).

Visiting Astronomers: Jack Gallimore and Michele Thornley, both of Bucknell University



**SOS Awards:** The SOS committee recommended funding a total of **\$106,500 to 4 of the 8 proposals submitted (only 7 of which were allocated observing time and considered for SOS funding)** this period. They are as follows: GBT10C-063, supervisor, Jim Jackson, student Susanna Finn, Boston University, for **\$20,000**; GBT10C-064, supervisor Mark Morris, student Elisabeth Mills, University of California, Los Angeles, for **\$35,000**; VLA10C-173, supervisor Lisa Young, student Kristina Nyland, New Mexico Tech, **\$19,500**; VLA10C-225, supervisor Min Yun, student Hansung Gim, University of Massachusetts, **\$32,000** 

Information on the SOS Program can be found at <u>http://science.nrao.edu/opportunities/sos.shtml</u>.

**Observatory Library:** The three NRAO workshop titles available electronically are as follows: **I.** Low frequency variability of extragalactic radio sources : proceedings of a workshop held at the National Radio Astronomy Observatory, Green Bank, West Virginia, April 21-22, 1982

**2.** Serendipitous discoveries in radio astronomy : proceedings of a workshop held at the National Radio Astronomy Observatory, Green Bank, West Virginia on May 4, 5, 6 1983

**3.** The search for extraterrestrial intelligence : proceedings of an NRAO workshop held at the National Radio Astronomy Observatory, Green Bank, West Virginia, May 20, 21, 22, 1985 / edited by K.I. Keller



**Archives:** Papers of *Ronald N. Bracewell*, the Papers of *David S. Heeschen*, and materials from the Director's Office. *Dr. Woodruff T. Sullivan III's* audio tapes are primarily interviews made over the period 1971-1988 with approximately 260 radio astronomers, including many pioneers in the field. About 20% of the tapes are of meeting talks related to radio astronomy history, and they will also be digitized. With the help of Josh Malone, processes and protocols for digitizing these fragile tapes have been developed, and the first several tapes were digitized before Dr. Sullivan's two-day visit to review the digitization done, discuss details of processing, and plan for future shipments of paper materials in his collection. In consultation with American Institute of Physics, Center for the History of Physics, we have developed procedures for obtaining the necessary oral interview donor agreements from interviewees or their heirs.

Finding aids for the Archives collection, and the Archives online catalog, are linked from the NRAO Archives home page, <u>http://www.nrao.edu/archives/</u>.





CCE: ALMA/NRAO High Availability servers now in place running the Helpdesk and Plone web-based Content Management System and User Portal including user authentication to JAO user database

Successfully completed Q4 integration test against requirements for ALMA observer supporting helpdesk.

NRAO Plone installed and configured to replicate the style and structure of www.nrao.edu, which will allow for workflow control over online document and content publication.

Improved user password security and digital signature certificate authority now running within NRAO

Network: Transitioned Long distance, International, Calling cards and voice/web Conferencing to the new Networx contract. Only the Brewster WA VLBA site remains from WAN upgraded in October.

Ongoing commitment to upgrade ALL major sites to **IGigabit/Second** to facilitate data access: Complete for C'ville and Socorro; WV State awarded \$8M million in economic stimulus funds to deliver 10Gbit/sec from Green Bank to WVU and on to the North American Research Networks: due Q4 FY 2011

A contract has been finalized with AURA for IGigabit/Second link from ALMA SCO to NA-ASC with connectivity in QI FYII

ARRA funds have been used to successfully replace the PBX and Voicemail system used throughout the Green Bank WV site.

Digital Infrastructure: **ALMA archive servers, User Portal and Oracle Database** now running in Charlottesville with initial test data now on-line.

The Next Generation **Science Data Archive** System servers now storing and replicating production EVLA data in Socorro leveraging high speed Lustre file system for staging data for CASA

Access to 200 TeraBytes of Archive storage and 100,000 hours of computer time being leveraged from NSF TeraGrid to support the re-use of the 350MHz GBT pulsar survey data.

Security: **IT Risk Assessment** review was successfully completed with Cherry, Bekaert & Holland, L.L.P. and a HIPAA review with Mercer. Recommendations to be executed throughout FY11

No **production** downtime from security vulnerabilities occurred this quarter.









In the Contracts and Procurement area in Q4, continue to implement and evaluation additional recommendations made in the Contractor Procurement System Review Report. New administrative requirements, and audits, are increasing Buyer's workload resulting in overtime and increasing the GB Buyer from .5 to .75FTE. The Buyers completed the year by issuing 6.077 purchase orders for approximately \$29,850,000.

Significant time and energy used to support the continuous, and increasing number, of audits. ARRA focus continues. All impacting existing resources and ability to process PR's and conduct the procurement process in a thorough and complete manner.

A procurement review was conducted in Chile for QI FY10 activity. Q2-Q4 scheduled for December 2010. Due to AUI's directive that's changed the travel budgeting/charging methodology, the NA ALMA Business Manager had performed the QI review in lieu of C&P, thereby reducing risks that could result if a review wasn't conducted.

ARRA expenditures continue to be a major focus, and soon ARRA audits. Working with NSF in Q4 to re-align remaining ARRA funds to new projects.

The CAP manuals have been escalated to a higher priority level.

There was an initiative for FY2010 to establish an expanded and user-friendly web presence in which internal and external customers will be able to obtain standard procurement forms, terms and conditions, representations and certifications, and proposal materials. The website is at risk and will not be updated as planned. Due the number of audits and re-prioritizations, the website update is being moved to FY2011.



**Edgemont Road Facility: Quarterly Building Maintenance Inspections completed.** Fire Extinguisher Inspections completed. Fire main and alarm maintenance completed in Stone Hall. Alarm tests successful. All routine. **Executive Safety meeting**-Travelled to OSF on Saturday after hearing of the fatality. The review was refocused into a meeting to support the investigation & rules changes.

**NTC** Now has an internally operated Fire Extinguisher management program. Also a volunteer Safety Coordinator to ensure that we take full advantage of internal resources before we use outside resources (\$\$\$).



**New Mexico Facilities-Outreach activities-**NMIMT Emergency Coordinator: Surveyed campus flood hazards and their mitigation as they affect NRAO employees and discussed various security issues including earthquake and "active shooter" awareness; NMIMT PD Chief: Discussed various security concerns and advised him of DSOC employee schedules to enable his department to provide additional security; Socorro County Sheriff: Discussed VLA site security issues and discussed DSOC & ROB security issues. **Blue Laser Research**-advised against it because of unreasonably high probability of inadvertent injury in the proposed application. **DSOC Safety walk through**-to educate in various safety issues: extension cord use, egress & ingress clearances. **Environmental/Green Initiatives:** transported about 120# of used batteries to VLA holding facility and transported about 15 gallons of used isopropyl alcohol to the VLA used oil facility. **Safety Training**-developed and implemented a process to distribute the Weekly Safety Topics to all of NRAO facilities (Safety Blog); developed short PowerPoint safety overview for visiting scientists (JPL), addressed five pertinent safety issues: elevation hazards associated with going from sea level to 7000°, low ambient humidity, poisonous reptiles, storm event safety, and personal security and sent to the Jet Propulsion Laboratory who will use it to supplement their own training; new DSOC-based employee safety orientation for five employees.



**New Mexico VLA Site: Meetings and Training Sessions**-training sessions offered in Lockout/Tagout and Cranes, Hoist & Slings and Slip, Trip and Fall awareness, as well as Understanding Safety at work; new employee safety orientations. The SO attended weekly coordination meetings, monthly safety committee meetings and ES&S coordination meeting. The EMS continued having monthly meeting time permitting.

**Ongoing Green actives:** continue with 2,960 Lbs of Lead-acid Batteries with Wise Recycling, 2,161 gallons of dyed diesel and oil with Mesa Environmental, 350 gallons of antifreeze, 300 gallons of heavy grease, and 200 gallons of machining oil. **Closed out final item for NMEPA PSTB NOV #3882**. The last item of the NOV 3882 was closed out with the replacement of the 25K gallon diesel tank. The tank was replaced with a 15k gallon doubled walled tank. New pump and leak detector system was also installed. **Outreach activities**-assisted Datil Fire Department with required NFPA hoses testing. SO assisted Magdalena Fire and EMS with ambulance and fire truck inspection and responded under mutual aid agreement. SO provided and participated with Magdalena ISD rock club, rocket building projects, SO provided tour for NRAO visitors under NRAO EPO. **Monthly and Quarterly Inspection and testing**-conducted monthly fire extinguisher inspection on all building fire extinguishers; conducted quarterly extinguisher inspection on all antenna fire extinguishers. EPA monthly generator logs completed and up-to-date. Biannual generator EPA air quality report completed and submitted. Monthly fuel tank inspection conducted and filed. NFPA 25 inspection and testing completed. **EMS refresher:** SO attended EMS refresher for license renewal.



**Green Bank Facilities: Recycling** efforts continue with 500 gallon of used oil recycled this quarter. **Site sewer water treatment system** continues within limits performance for the third successive year. Annual **site OSHA inspection** was conducted in July with no major issues to report. **Training** was conducted in the area of Respiratory Protection refresher and new employee safety orientations.

VLBA Facilities: All sites updated with first aid kits and/or supplies.



MRI-R2 VLBA Sensitivity period of performance: April 1, 2010 to March 8, 2011 Research Opportunities in Space and Earth Science: April 15, 2010 to April 14, 2011



Recognition in Diversity Careers magazine resulted in # diversity applicants seeking employment at NRAO.



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NRAO Operations (less EVLA) FY 2010 new funding allocation is \$43,144k. Total available funding including prior year commitments and carryover totals \$48,879k. Total spending for FY 2010 is \$45,982k with the remaining \$2,897k to cover anticipated FY 2011 and FY 2012 funding shortfalls.