# NATIONAL RADIO ASTRONOMY OBSERVATORY CHARLOTTESVILLE, VA

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# NOISE PARAMETERS OF NRAO 1.5 GHz GASFET AMPLIFIERS

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#### I. Introduction

It is common practice in the low-noise microwave amplifier field to only partially specify the amplifier noise performance. The noise parameter which is usually reported is  $T_A$ , the noise temperature of the amplifier when driven from a particular source impedance, which is usually 50 ohms for coaxial input lines. In order to analyze the performance of a system in which the amplifier is driven from some other source impedance, a set of four noise parameters of the amplifier must be known. There are many sets of noise parameters; six sets and the transformations between them are described in [1]. The most common set of parameters are the minimum (vs. source impedance) noise temperature,  $T_{min}$ , the impedance,  $Z_{opt}$ , which gives this minimum, and a sensitivity parameter,  $G_n$ . The particular set of parameters most convenient for the work described here is the noise wave model described by Penfield [2] and Meys [3] and shown in Figure 1c.

The noise parameters, with the exception of a phase angle, \* for the 1.5 GHz low-noise amplifier reported in [4] will be described in this report. Approximately

The missing phase angle is the phase of the wave correlation coefficient which is used to determine the phase of the optimum source reflection coefficient,  $\Gamma_{opt}$ . Data sufficient to determine this angle was collected but the accuracy of the data is not adequate to determine the angle accurately, especially when  $|\Gamma_{opt}|$  is small. The angle may be the subject of some future work but was not considered as a necessity for our present study.

ninety of these amplifiers are under construction by NRAO for use as cooled receiver front-ends and as I.F. amplifiers for cooled Schottky-diode and superconducting tunnel junction millimeter-wave mixer receivers. Questions regarding the effect of mixer-I.F. mismatch have arisen and this is the prime motivation for the measurements. It is also very useful to know the difference between  $T_A$  and  $T_{min}$  - i.e., how much improvement in noise temperature can be obtained by redesign of the input network.

#### II. Measurement Method

The measurement configuration and analysis models are shown in Figure 1. When  $\Gamma_s$  is a 50-ohm resistor at known temperature, the Apple computer program NOISE1 tabulates and plots  $T_A$ , the amplifier noise temperature at  $\Gamma_s = 0$ , using values of noise diode excess temperature and source noise temperature with diode off,  $T_{OFF}$ , all referred to the amplifier input connector at 15K. These calibration values have been determined by applying hot and cold noise temperature standards [5] at reference plane A and measuring the A to B loss.

After the measurement of  $T_A$ , the 50-ohm resistor is replaced by a sliding short,  $T_{OFF}$  is changed in the program, and noise measurements are made for several positions of the sliding short. An appropriate value of  $T_{OFF}$  is the noise temperature delivered to a noiseless 50 ohm load replacing the amplifier input; i.e.,  $T_p(1 - |\Gamma_g|^2)$  where  $T_p$  is the physical temperature of the lossy part of  $\Gamma_s$ . If this is done, NOISEl tabulates and plots a quantity which we will call the amplifier noise wave temperature,  $T'_n = T_n(1 - |\Gamma_g|^2)$ , where  $T_n$  is the amplifier noise temperature. Both  $T'_n$  and  $T_n$  are functions of  $\Gamma_s$  but as  $|\Gamma_s| + 1$ ,  $T'_n$  is well behaved and  $T_n \neq \infty$ . Typical plots of  $T'_n$  versus frequency for three different amplifiers and two different short positions as produced by NOISE1 are shown in Figure 2.

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Fig. 1: The measurement physical configuration is shown in (a) where  $\Gamma_s$  is either a 50-ohm resistor at 15K or a sliding short at 300K. The Apple-computer output is then the noise-wave temperature,  $T'_n$ , shown in (b). This is then used to find parameters of the noise-wave model shown in (c).



Noise wave temperature,  $T'_n$ , at sliding short positions separated 1 5.08 cm for 3 different amplifiers at 15K. The main point is that the amplifiers are very similar; detailed analysis was performed only on one amplifier #169 (middle plot).

The noise wave temperature,  $T'_n$ , is only of intermediate interest. It is easily measured and is related to the noise wave model of the amplifier shown in Figure 1c by,

$$T'_{n} = T_{A} + T_{B} |\Gamma_{s}|^{2} + 2 \sqrt{T_{A}T_{B}} \operatorname{Re}(\rho\Gamma_{s})$$
(1)

where  $T_A$  and  $T_B$  are the ingoing and outgoing noise wave temperatures and  $\rho$  is the complex correlation coefficient between  $T_A$  and  $T_B$ . These noise parameters are properties of the amplifier (i.e., independent of  $\Gamma_S$ ) and can be transformed to other parameters.

The value of  $T_{OFF} = T_p(1 - |\Gamma_s|^2)$  is difficult to determine accurately. An initial estimate was first used in the NOISE1 program and the resulting  $T_n^i$  was then corrected to give a value of  $T_B$  which agreed with a separate direct measurement of  $T_B$ . This direct measurement was performed by connecting the amplifier <u>input</u> port to a receiver which has been calibrated with hot and cold termination. The outgoing noise from the amplifier input,  $T_{OUT}$  is approximately  $T_B$  for a low value of amplifier input reflection coefficient,  $|\Gamma_n|$  (which is < .1 at band center) and incident outgoing receiver noise,  $T_{ISO}$ , comparable to  $T_B$ . The results of this measurement for an amplifier at 300K and 15K are shown in Figure 3. The resulting value of  $T_{OFF}$  is 26K. Assuming  $T_p = 300$ K, this gives  $|\Gamma_s| = .955$  or .393 dB return loss (.196 dB one-way loss) for a Maury 1929-2 sliding short, a 20 dB coupler, and a dewar transition.

It should be remarked that some difficulty was experienced with three different Maury 1929 sliding shorts which were utilized. Two of the shorts gave erratic results due to dirty contacts and loose bits of metal in the short; much improvement was noted after disassembly and cleaning. The interior surfaces of the coaxial lines appear to be brass and better results may be obtained if they were gold-plated. A 4K difference in noise temperature was measured between short positions spaced  $\lambda/2$  apart at 1.5 GHz; this gives a one-way loss of .003 dB/cm. The final data taken on the cooled amplifier was repeated wit two different sliding shorts with fairly good agreement; i.e., for  $|\Gamma_{opt}|$  results were within .05 from 1.3 to 1.6 GHz and within .20 at all frequencies.

#### III. Results

An Apple sub-program, SS2 PLOT, used as part of a LADDER main program (which provides plotting and utility functions), was written to analyze  $T_n'(f)$ data versus sliding short position; the program is listed in Appendix II. Equation (1), written for four values of  $\Gamma_s$  is inverted to give  $T_A$ ,  $T_B$ , and  $\rho$ magnitude and phase. The measurement at  $\Gamma_s = 0$  gives directly  $T_A$  and measurements at  $|\Gamma_s| = .955$  and three values of phase can be fitted to a sinusoidal function as shown in Figure 4. The resulting three noise parameters for amplifier #169 at 15K and 300K and frequencies of 1.2 to 1.9 GHz are given in the first three columns of Table I. The last three columns give the quantities,  $T_{min}$ ,  $T_D$ , and  $\Gamma_{opt}$ , which are convenient for describing the amplifier noise temperature,  $T_n$ , vs. source reflection coefficient,  $\Gamma_s$ , as:

$$\Gamma_{n} = T_{min} + \frac{T_{D} |\Gamma_{s} - \Gamma_{opt}|^{2}}{1 - |\Gamma_{s}|^{2}}$$
(2)

where

$$T_{\min} = \frac{T_{A} - T_{B}}{2} + \sqrt{\left(\frac{T_{A} + T_{B}}{2}\right)^{2} - T_{A}T_{B}|\rho|^{2}}$$
(3)

$$T_{\rm D} = T_{\rm min} + T_{\rm B} \tag{4}$$

$$|\Gamma_{opt}| = |\rho| \cdot \frac{\sqrt{T_A T_B}}{T_D}$$
(5)

Note that when  $|\rho| = 0$  then  $|\Gamma_{opt}| = 0$  and  $T_{min} = T_A$ .

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Fig. 3: Noise out of amplifier input terminals with amplifier at 300K (top two curves) and 15K.  $T_{\rm ISO}$  is the noise temperature out of the measuring receiver, incident and partially reflected by the amplifi At 1.5 GHz the amplifier outgoing noise wave,  $T_{\rm B}$ , is found to be 43K with the amplifier at 300K and 8K with the amplifier at 15K.



MAURY 1929A SHORT, AMP #169

Fig. 4: Noise wave temperature,  $T'_n$ , of amplifier #169 at 15K and 1.5 GHz as a function of sliding short position (inches). The dotted line is an exact fit to the left 3 points. The right 2 points show effect of additional loss as the short was extended.

Examination of (2) shows that, for a given value of  $|\Gamma_s|$ , the minimum of  $T_n$  with respect to phase of  $\Gamma_s$ , occurs at  $|\Gamma_s - \Gamma_{opt}| = ||\Gamma_s| - |\Gamma_{opt}||$  and maximum at  $|\Gamma_s| + |\Gamma_{opt}|$ . Thus the range of  $T_n$  can be plotted vs.  $|\Gamma_s|$  for typical midband values of noise parameters. This is shown in Figure 5 along with the  $T_n$  vs.  $|\Gamma_s|$  curve assuming an ideal isolator between source and amplifier. Note that at 300K the isolator always increases  $T_n$  but at 15K this is not true.



Ranges of noise temperature vs. source reflection coefficient magnitude with and without isolators for amplifiers at 300K ( and 15K. Typical values of  $|\Gamma_{opt}| = .1$ ,  $T_{min} = 60$ ,  $T_D = 120$  300K and  $|\Gamma_{opt}| = .3$ ,  $T_{min} = 8$ ,  $T_D = 18$  at 15K are assumed.

ſ		N	oise Wave	S	See	Eqs. (2)-(5	)
	F GHZ	<sup>Т</sup> А °К	т <sub>в</sub> °к	p	T min °K	<sup>т</sup> р °к	r <sub>opt</sub>
	1.2	22.1	20.8	.75	15.0	35.7	.45
	1.3	14.1	12.9	.70	10.3	23.2	.40
	1.4	9.8	6.8	.64	7.9	14.7	.36
	1.5	9.8	8.9	.46	8.8	17.7	.24
	1.6	9.1	8.7	.59	7.4	16.0	.33
	1.7	10.0	7.1	.47	9.0	16.1	.25
	1.8	8.7	7.6	.10	8.7	16.2	.05
	1.9	15.7	8.8	.76	11.9	20.7	.43
+							
	1.2	109	72	.68	85	157	.38
	1.3	81	61	.43	74	135	.23
	1.4	61	51	.21	60	110	.11
	1.5	66	47	.11	66	113	.05
	1.6	63	52	.18	62	114	.09
	1.7	65	56	.08	65	121	.04
	1.8	68	63	.19	67	130	.10
	1.9	106	63	.59	91	153	.31

# TABLE I - Noise Parameters of Amplifier #169 at 15K (top) and 300K (bottom)

#### References

- [1] D. L. Fenstermacher, "A Computer-Aided Analysis Routine Including Optimization for Microwave Circuits and Their Noise," NRAO Electronics Division Internal Report No. 217, July 1981.
- [2] P. L. Perifield, "Wave Representation of Amplifier Noise," <u>IRE Trans</u>. <u>Circuit Theory</u>, vol. CT-9, pp. 84-86, March 1962. Also reprinted <u>in Low-Noise Microwave Transistors and Amplifiers</u>, edited by H. Fukui, Wiley, 1981.
- [3] R. P. Meys, "A Wave Approach to the Noise Properties of Linear Microwave Devices," <u>IEEE Trans. Mic. Th. and Tech.</u>, vol. MTT-26, pp. 34-37, Jan. 1978. Also reprinted in book described in [2].
- [4] S. Weinreb, D. Fenstermacher, and R. Harris, "Ultra Low-Noise, 1.2-1.7 GHz Cooled GASFET Amplifiers," <u>IEEE Trans. Mic. Th. and Tech.</u>, vol. MTT-30 no. 6, June 1982. Also published as NRAO Electronics Division Interna Report. No. 220, revised May 1982.
- [5] S. Weinreb, "Comparison of Three Liquid Nitrogen Noise Standards at 1.5 and 4.75 GHz," NRAO Electronics Division Technical Note No. 101, September 11, 1981.

### APPENDIX I - Quick Use of Sliding Short to Determine Noise Parameters

A setup such as Figure 1a, with either a computer or noise figure meter to indicate noise temperature, can be used to quickly give noise parameters less the phase angle. With  $\Gamma_s$  a matched load the noise temperature measured in the usual way is  $T_A$ . The matched load is then replaced by a sliding short and  $T_{OFF}$  within the computer is replaced by  $300(1 - |\Gamma_{ss}|^2)$  where  $|\Gamma_{ss}|$  applies to the sliding short; alternatively  $300 \times |\Gamma_{ss}|^2$  can be substracted from the indicated  $T_n$ . The sliding short is then moved to find the high,  $T_H$ , and low,  $T_L$ , extremes.  $T_B$  and  $|\rho|$  are then computed as:

$$T_{B} = [(T_{H} + T_{L})/2 - T_{A}]/|\Gamma_{ss}|^{2}$$
(6)

$$|\rho| = \frac{T_{H} - T_{L}}{4 \times |\Gamma_{ss}| \sqrt{T_{A}T_{B}}}$$
(7)

Other parameters are then computed by equations (3)-(5).

APPENDIX II - SS2 Plot Subprogram

```
100
     REM
            SLIDING SHORT ANALYSIS, DECEMBER 21, 1980
101
     UTAB (21)
     INPUT "L1,L2,L3,L4,AND L5 (INCHES) ? ";L1,L2,L3,L4,L5
105
106 \ 60 = .40
110 \text{ GS} = \text{EXP} (-.11513 * \text{GD}); \text{REM}
                                          COUPLER+SLIDING SHORT LOSS
112 GOTO 115
113 F = 1.5:T0 = 10.4:T1 = 46.6:T2 = 40.7:T3 = 28.6:T4 = 38.2:T5
= 48.3: GOTO 117
115 INPUT "F(GHZ),T0,T1,T2,T3,T4,T5? ";F,T0,T1,T2,T3,T4,T5
117 XL = - .50:XH = 4.5:YL = 000:YH = 200: GOSUB 10200
118 XP = L1:YP = T1: GOSUB 10400
119 XP = L2: YP = T2: GOSUB 10400
                                          PLOT
                                          SUBROUTINE
120 \text{ XP} = \text{L3:YP} = \text{T3: GOSUB} 10400
121 XP = L4:YP = T4: GOSUB 10400
122 XP = L5: YP = T5: GOSUB 10400
               SET UP INTERMEDIATE PARAMETERS
124 REM
125 BB = 53.19776E - 2 * F:PI = 3.14159
130 \text{ A1} = \text{BB} * (\text{L2} - \text{L1}); \text{A2} = \text{BB} * (\text{L3} - \text{L1})
135 \text{ A3} = \text{BB} * (\text{L2} + \text{L1}); \text{A4} = \text{BB} * (\text{L3} + \text{L1})
140 GZ = (T2 - T1) * SIN (A2) / ((T3 - T1) * SIN (A1))
145 \text{ NZ} = 4: JM = 10
150 EM = 1E10:P1 = 0:P2 = PI:P3 = P2 / NZ
             SOLVE FOR P BY ITERATION
155
    REM
160
    FOR J = 1 TO JM
165
    FOR P = P1 TO P2 STEP P3
170 EZ = 6Z * SIN (A4 + P) - SIN (A3 + P)
    IF ABS (EZ) < EM THEN EM = ABS (EZ):PM = P:EY = EZ
175
180
    NEXT P
185 P1 = PM - P3:P2 = PM + P3:P3 = P3 / NZ
190
     NEXT J
195
     REM
            P=PM IS NOW KNOWN. FIND TP AND TU NEXT
200 TP = (T1 - T2) / (2 * SIN (A1) * SIN (A3 + PM))
    IF TP < 0 THEN TP = - TP:PM = PM - PI
205
210 TU = T1 - TP * COS (2 * BB * F * L(1) + PM)
215
    REM
             FIND NOISE WAVE PARAMETERS, TA, TB, AND RHO.
220 \text{ TA} = \text{T0}
225 TB = (TV - TA - 0 + (1 - 6S \wedge 2)) / 6S \wedge 2
    IF TB < 0 THEN TB = .001
230
235 RH = TP / (2 * SQR (TA * TB) * GS)
             FIND NOISE PARAMETERS TMIN, GOPT, GARG, AND TD
240
    REM
245 TC = RH \star SQR (TA \star TB)
250 TD = 0.5 * (TA + TB + _ SQR ((TA + TB) ^ 2 - 4 * TC ^ 2))
255 TMIN = TD - TB:GOPT = TC / TD:GARG = 3.14159 - PM
260 RH = INT (1000 * RH + .5) / 1000
    DEF FN R1(X) = INT (10 + X + .5) \times 10
265
     DEF FN R3(X) = INT (1000 * X + .5) / 1000
270
275
     PRINT
     PRINT F;"MHZ TMIN="; FN R1(TMIN);" TA="; FN R1(TA);" TB=";
280
 FN R1(TB);" TD="; FN R1(TD);" GAMMA OPT="; FN R3(GOPT);","; FN R3
 RHO="; FN R3(RH);","; FN R3(PM);" TU="; FN R1(TU);
300 DL = 16 / 250
    FOR XP = - .5 TO 4.5 STEP OL
310
320 YP = TU + TP * COS (2 * BB * XP + PM)
330 XT = D8 * XP + D9:YT = D6 * YP + D7
    IF XT < X8 THEN XT = X8
332
333
     IF XT > X9 THEN XT = X9
     IF YT < Y9 THEN YT = Y9
334
     IF YT > Y8 THEN YT = Y8
335
```

```
336 HPLOT XT,YT
340 YT = D6 * T0 + D7: HPLOT XT,YT
345 NEXT XP
350
    GET ZZ$: GOTO 115
800 REM IN LIMITS US REFLECTION COEF.
805 \text{ TM} = 8:\text{TD} = 23:\text{P} = .0
810 XL = 0:XH = 1:YL = 0:YH = 50: GOSUB 10200
820 FOR S = 0 TO .95 STEP .05
825 TX = TM + TD * (S + P) * (S + P) / (1 - S * S)
830 TN = TM + TD \star (S - P) \star (S - P) / (1 - S \star S)
835 RL = 8.68 * LOG (S + .0001)
840 WI = 8:DI = 2
842 X = S: GOSUB 11000
                                FORMAT
844 X = RL: GOSUB 11000
                                SUGROUTINE
846 X = TX: GOSUB 11000
                            1
848 X = TN: GOSUB 11000
850 PRINT
855 XP = S:YP = TN: GOSUB 10400
860 XP = S:YP = TX: GOSUB 10400
865 NEXT S: END
```