

North American ALMA Science Center



ALMA Correlator Capabilities for Early Science (Cycle 0)

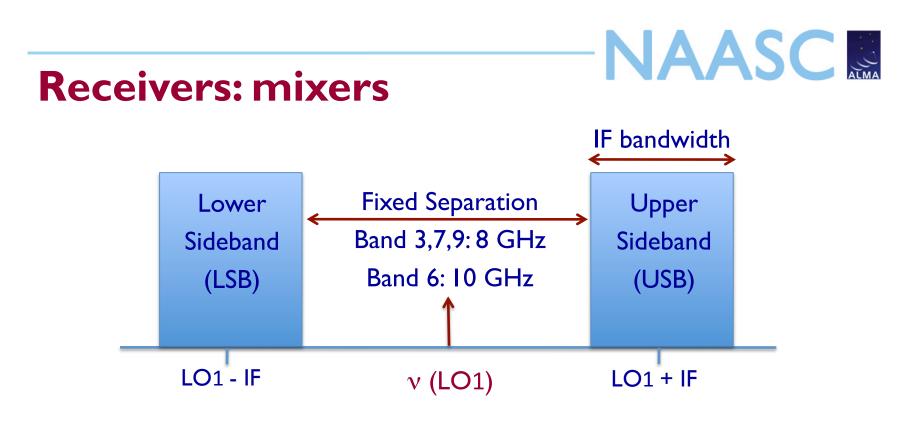
NAASC Memo #105

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ABSTRACT

The following slides summarize the Cycle 0 capabilities of the ALMA correlator. They compile the basic information from the ALMA Science Portal and the Cycle 0 Technical Handbook, and include some correlator set-up examples with screenshots from the ALMA Observing Tool. These slides are meant to be included in the NA presentations for ALMA Community Days Events.

A power point version of the slides can be found at https://sites.google.com/site/almacommunityoutreach/slides-folder/alma-capabilities.



- The first Local Oscillator (LO1) can be tuned at different frequencies
- The Sky Frequencies will be: $v_{sky} = v_{LO1} v_{IF}$ (LSB)

$$v_{sky} = v_{LO1} + v_{IF}$$
(USB)

 Data will be collected setting-up different spectral windows within one or both sidebands



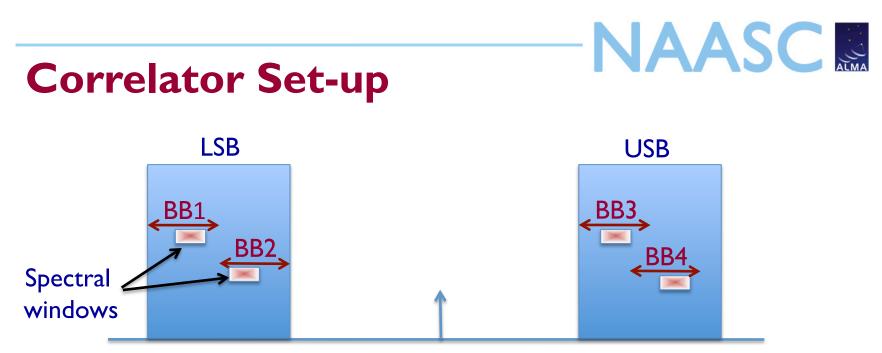


Cycle 0 Receivers

Band	Freq Range (GHz)	Wavelength range (mm)	Receiver type	IF range (GHz)	Inst. IF bandwidth (GHz)
3	84-116	3.6-2.6	2SB	4-8	8
6	211-275	1.4-1.1	2SB	5-10	8
7	275-373	1.1-0.8	2SB	4-8	8
9	602-720	0.5-0.4	DSB	4-12	8*

- Dual side-band (2SB): two separated sidebands available simultaneously
- Double side-band (DSB): LSB and USB are super-imposed out from the receiver but can be separated in later processing.
 - * Cycle 0: In Band 9, only one sideband per spectral window will be correlated. In future cycles both sidebands will be simultaneously separated and correlated.





LO1

- Up to 4 basebands available which can be moved within the sideband width; spectral windows can be moved within the baseband (2 GHz wide)
- Setup limits: Edges of the baseband cannot lie outside the IF range & edges of the spectral window cannot lie outside the baseband
 - * Cycle 0: only one spectral window per baseband & all spectral windows with the same configuration (bandwidth and spectral resolution)





Correlator Set-up

- Correlator modes:
 - Time division mode (TDM): low-spectral resolution \rightarrow continuum observations and sources with very broad spectral lines.
 - Frequency division mode (FDM): high-spectral resolution mode
 - * Cycle 0: One TDM and 6 FDM set-ups available
- Dual or Single Polarizations can be processed:
 - Dual Polarization: separate spectra obtained for each linear polarization \rightarrow can be combined to improve sensitivity
 - Single Polarization: only a single polarization is analyzed
 - \rightarrow poorer sensitivity but provides twice as many channels





Cycle 0 Correlator Modes

Dual Polarization

Single Polarization

Bandwidth (MHz)	Channel spacing (MHz)	Number of channels	Bandwidth (MHz)	Channel spacing (MHz)	Number of channels
2000	15.6	128*	2000	7.8	256*
58.6	0.0153	3840	58.6	0.0076	7680
117	0.0305	3840	117	0.0153	7680
234	0.061	3840	234	0.0305	7680
469	0.122	3840	469	0.061	7680
938	0.244	3840	938	0.122	7680
1875	0.488	3840	1875	0.224	7680

* Correlator mode for **continuum** observations (TDM), effective bandwidth is only 1875 MHz

In all cases, the effective resolution is twice the channel spacing

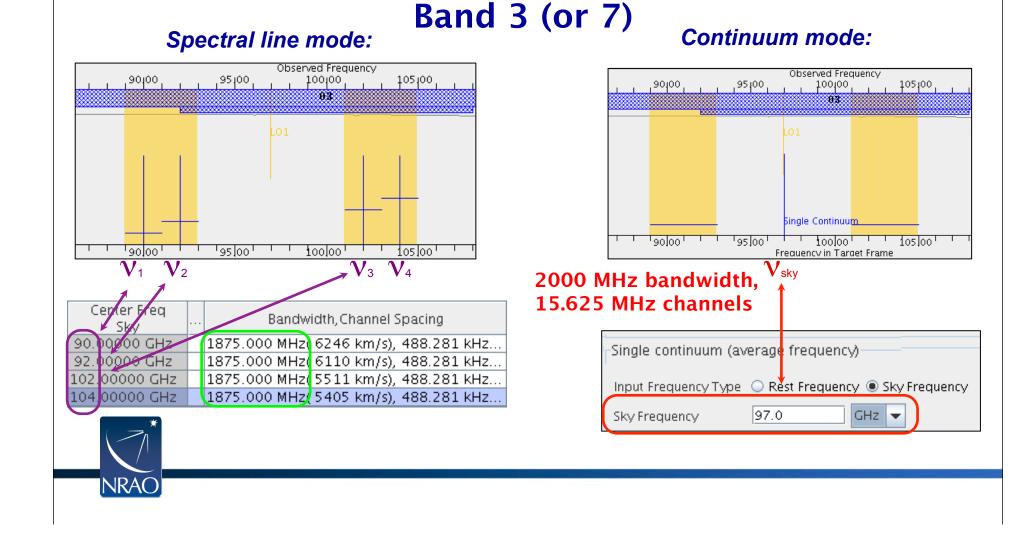


Correlator set-up:

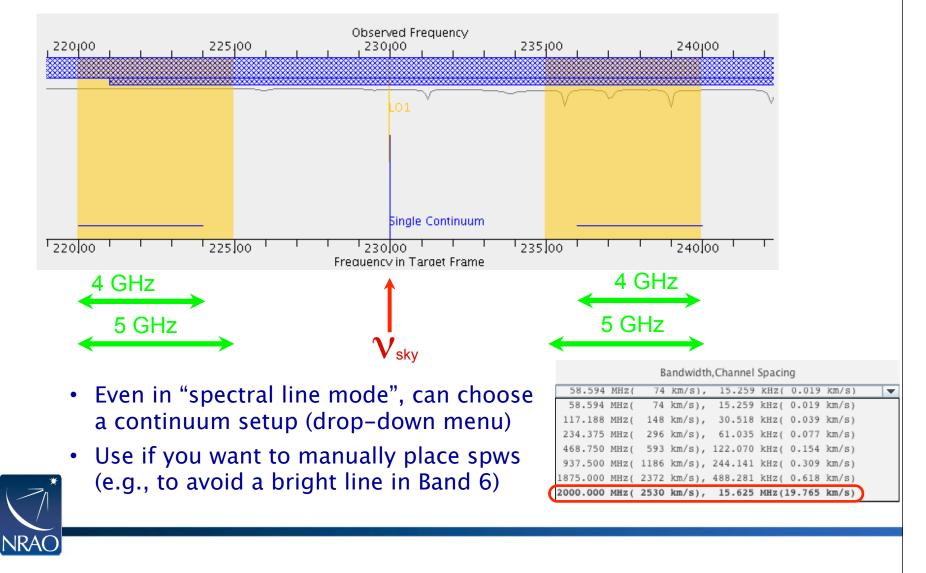
line vs. continuum

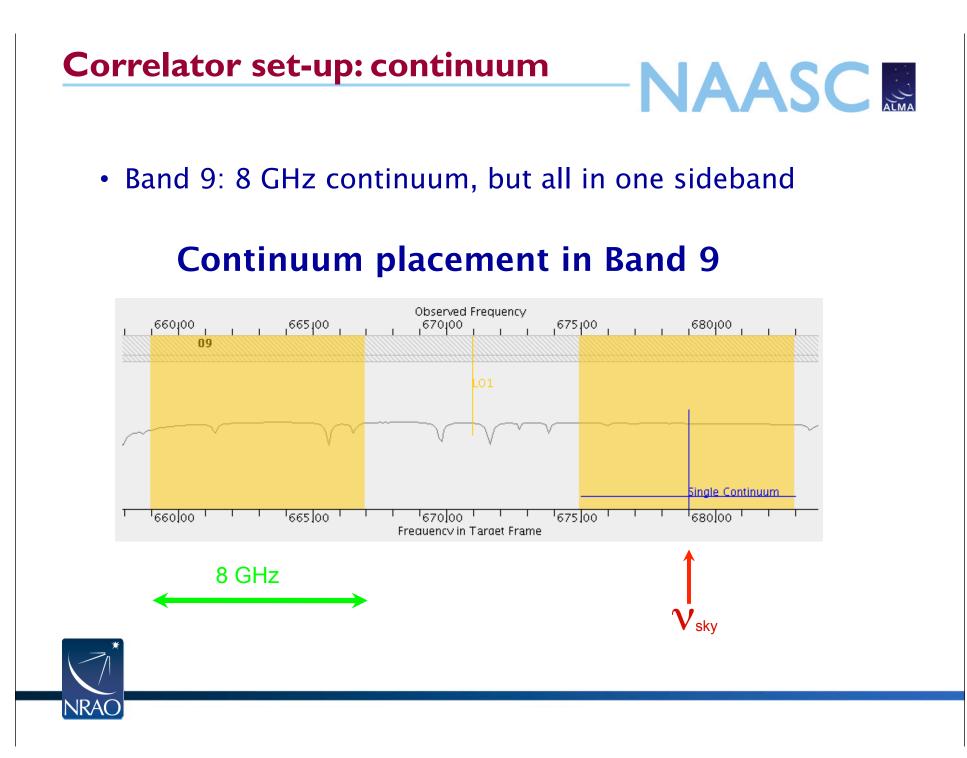
- NAASC

• "continuum mode": automatically place 4 spectral windows, with the largest bandwidth, across the sidebands



Standard continuum placement in Band 6





Correlator set-up: continuum



- Select "single continuum" in OT
- Single vs Dual polarization allowed

-Spectral Type			
		 Up to 4 spectral windows 	
	Spectral Type: Choose the type of spectra	$ \bigcirc$ More than 4 spectral windows	
	observation you wish to make	Single continuum (average frequency)	
		 Spectral scan 	
	Polarization Products desired	⊖ SINGLE-X	

- Single polarization provides same total bandwidth, but twice the spectral resolution: lower sensitivity for averaged continuum!
- (Unless your science case is very special, you should always select Dual polarization for continuum observations.)





Survey of carbon isotopologues and other species in old planetary nebula

Observe molecular gas in NGC 7293 (Helix Nebula) in Band 7

Main species of interest

<u>Lower sideband (LSB)</u>

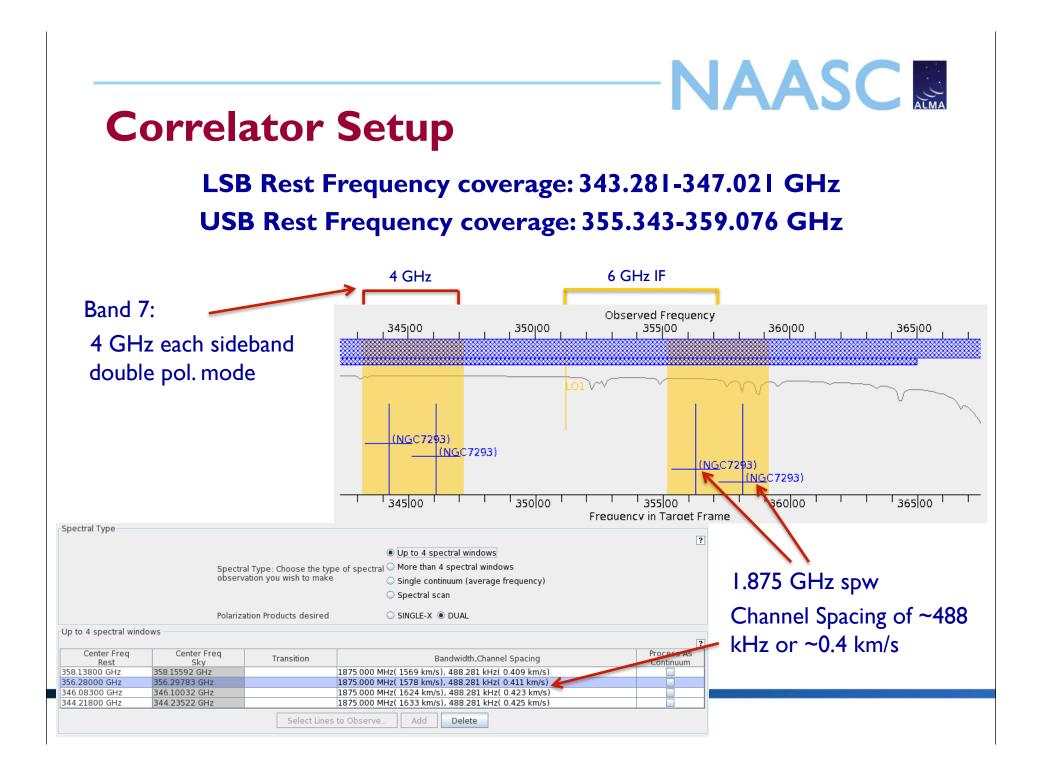
CO (J=3 \rightarrow 2) : 345.795 GHz H¹³CN (J=4 \rightarrow 3) : 345.339 GHz H¹³CO⁺ (J=4 \rightarrow 3): 346.998 GHz <u>Upper sideband (USB)</u>

HCO⁺ (J=4 \rightarrow 3): 356.734 GHz

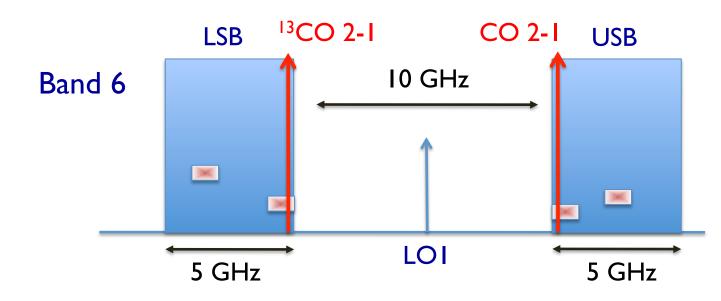
Other interesting molecular species present within the spectral setup

> CS, CS⁺, HCP, MgCCH, NaCN, MgH, SiH, MgNC, KC, AINC, SO, ³⁴SO





Example: ¹²CO and ¹³CO in Band 6

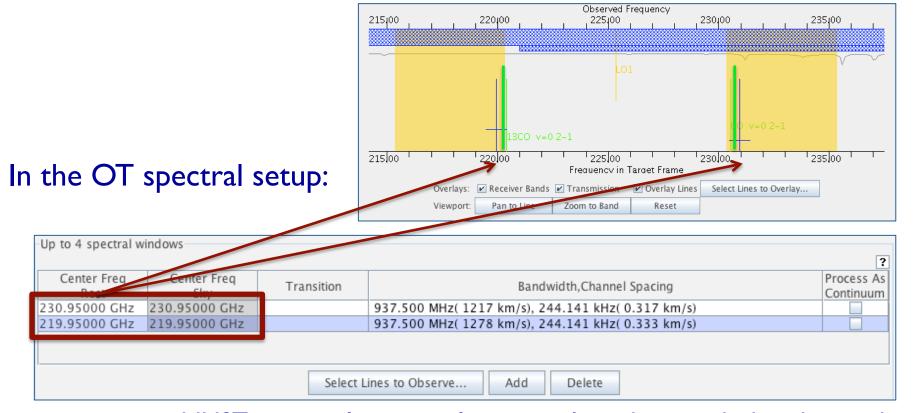


- 2 spectral windows, 0.9375 GHz wide, 0.3 km/s spectral resolution
- Can observe both ^{13}CO 2-1 (220.4 GHz) and CO 2-1 (230.5 GHz) only at low z
 - MUST set rest frequency for spectral windows such that the windows remain entirely within the sidebands, e.g. can't center on lines for wider spectral windows



can place 2 additional windows to observe CH_3OH , SO_2 , etc.

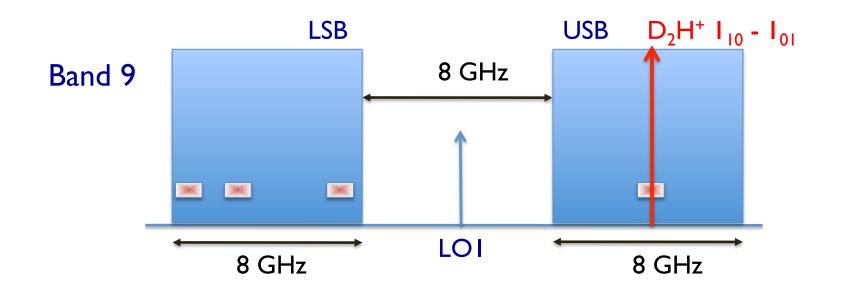
Example: ¹²CO and ¹³CO in Band 6



- MUST set rest frequency for spectral windows such that the windows remain entirely within the sidebands (i.e. not centered on lines)
 - can place 2 additional windows to observe CH_3OH , SO_2 , etc.

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Example: Spectral Lines in Band 9



- for Band 9, there is full flexibility in that each baseband can be connected to either one or the other sideband
- in Cycle 0, only one sideband per spectral window will be correlated
 - e.g. Observe D_2H^+ at 691.66 GHz with one spectral window



• can place 3 additional windows in USB or LSB