

Experimental Tests of General Relativity

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1) General Relativity is Einstein's theory of gravitation.

Unlike Newtonian gravitational theory, it is invariant under the Lorentz transformation, and ~~it~~, in a very beautiful way, accounts for the universality of the acceleration of freely falling bodies. Rather than describe gravity as a force, in General Relativity theory (GRT), freely falling particles are in natural motion in a curved space; gravity becomes a property of space. It is the curvature of space rather than a force which brings bodies together.

2) The curvature of space-time is obtained from the metric of space-time $g_{\mu\nu}$, and its derivatives.

3) For flat space $g_{\mu\nu} = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix} = \eta_{\mu\nu}$ the Minkowski metric tensor.

$$ds^2 = g_{\mu\nu} dx^\mu dx^\nu$$

$$= dx^0 dx^0 - dx^1 dx^1 - dx^2 dx^2 - dx^3 dx^3$$

$$= dt^2 - dx^2 - dy^2 - dz^2$$

4) Curvature is expressed by $R_{\mu\nu}$, the Ricci tensor and R , the curvature scalar, which are functions of $g_{\mu\nu}$ and its derivatives.

5) The Einstein field equations relate the curvature of space-time to the distribution of matter

$$R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = T_{\mu\nu} = \begin{pmatrix} \text{(Energy density)} & \text{(Momentum density)} \\ \text{---} & \text{---} \\ \text{---} & \text{(stress tensor)} \end{pmatrix}$$

$T_{\mu\nu}$ = stress-energy tensor.

6) GRT is only one of over 20 ^{relativistic} metric theories of gravity.

7) A metric theory :

- A) endows space with a metric
- B) tests bodies move along geodesics in the metric
- C) in a freely falling frame, special relativity holds

8) The many metric theories differ only in the way in which matter or other scalar, vector, or tensor fields ~~and~~ determine $g_{\mu\nu}$.

9) Example : IN GRT

$$\left(\begin{array}{l} \text{Matter + energy dist} \\ \rightarrow T_{\mu\nu} \end{array} \right) \Rightarrow \left(\begin{array}{l} \text{elements of} \\ g_{\mu\nu} \end{array} \right)$$

in scalar-tensor theory of Brans & Dicke

$$T_{\mu\nu} \Rightarrow (\text{scalar field, } \phi)$$

$$(T_{\mu\nu}, \phi) \Rightarrow (g_{\mu\nu})$$

10) To determine particle motion under various theories, find weak field limits and write very general metric with parameters determining influence of ^{grav.} energy, mass, rest mass, momentum, etc. on metric. Each theory of gravity will determine a value for these parameters.

11) Express motion of particles ~~interms~~ in terms of these parameters and experimentally find their values, thereby supporting some theories for a true theory of gravity

Example: simple parametrized post-Newtonian metric

(Eddington - Robertson)

$$g_{\mu\nu} = \begin{pmatrix} g_{00} & & & \\ & g_{11} & & 0 \\ & & g_{22} & \\ & & & g_{33} \end{pmatrix}$$

$$g_{00} = 1 - 2U + 2\beta U^2$$

$$g_{11} = g_{22} = g_{33} = -(1 + 2\gamma U)$$

parameters β, γ

β measures non-linearity of superposition of fields
 γ " " curvature of space

for Newtonian gravitation $\gamma = \beta = 0$
 GRT $\gamma = \beta = 1$

scalar-tensor theory (B.D.) $\gamma = \frac{1+\omega}{2+\omega}$ $\beta = 1$

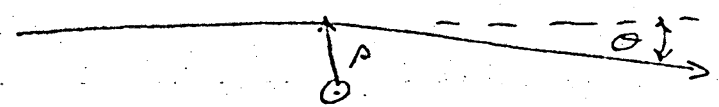
$\omega =$ coupling constant ~ 6
 $\therefore \gamma \sim -0.875$

13) Experiments.

light bending $\Delta\theta = \frac{1}{2}(1+\gamma) \frac{1.75}{\rho}$ $\rho = \frac{R}{R_0}$

predictions:

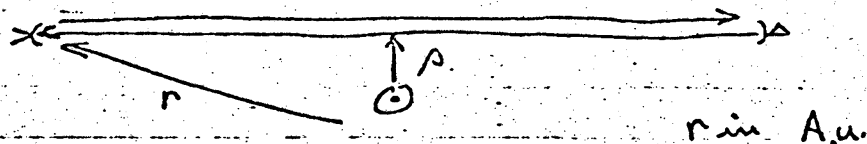
$\lambda = \frac{1}{2}(1+\gamma) = 1.00$ GRT
 $= 0.94$ B.D.
 ($\omega \sim 6$)



observations: radio measurements.

1969	$\lambda = 1.04 \pm .15$	JPL.	1970	} $\lambda = 0.94 \pm 0.06$	N.R.A.O.
	$\lambda = 1.01 \pm .12$	O.V.R.O.	1971		

14) time delay:



$$\Delta T = \lambda \left[250 - 20 \ln \frac{\rho^2}{r} \right] \text{ usec}$$

passive radar to ~~Earth~~ Mercury

$$\lambda = 1.02 \pm 0.05$$

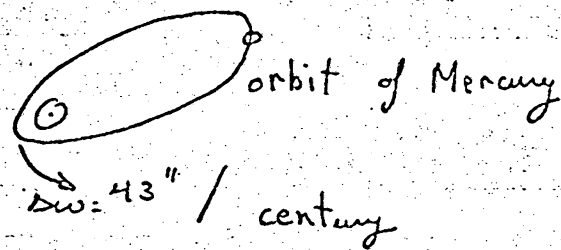
active radar to Mariner VI
VII

$$\lambda = 1.00 \pm 0.04$$

15) Perihelion advance:

$$\Delta \omega = 43 \left[\frac{1}{3} (2\gamma + 2 - \beta) \right] + 4 \left[\frac{J_2}{3 \cdot 10^{-5}} \right]$$

where J_2 = quadrupole moment of Sun.



$$\Delta \omega \text{ measured} = 43 \pm 0.4 \text{ arc sec / century}$$

$$\text{if } J_2 = 0$$

$$\frac{1}{3} (2\gamma + 2 - \beta) = 1$$

$$\therefore \gamma = 1.0$$

assume $\beta = 1$

GRT is right!

$$\text{if } J_2 = 3 \cdot 10^{-5}$$

$$-\frac{1}{3} (2\gamma + 2 - \beta) = \frac{39}{43}$$

$$\therefore \gamma = 0.86$$

B.D. is right!

measured $J_2 \sim 3 \cdot 10^{-5}$ but interpretation is still questioned.

TABLE 5.1. Metric Theories of Gravity and Their PPN Parametric Values

Theory and its adjustable parameters [†]	γ	β	α_1	α_2	α_3	ξ_1	ξ_2	ξ_3	ξ_4
1. General Relativity ^L	1	1	0	0	0	0	0	0	0
2. Scalar-Tensor Theories									
a. Bergmann-Wagoner-Nordvedt (ω, Λ) ^L	$\frac{1+\omega}{2+\omega}$	$1+\Lambda$	0	0	0	0	0	0	0
b. Dicke-Brans-Jordan (ω) ^L	$\frac{1+\omega}{2+\omega}$	1	0	0	0	0	0	0	0
3. Vector-Metric Theory (K) ^L	1	1	0	$\frac{K^2}{1+\frac{1}{2}K^2}$	0	0	0	0	0
4. Conformally Flat Theories									
a. Nordström-Einstein-Fokker ^L	-1	$\frac{1}{2}$	0	0	0	0	0	0	0
b. Ni's Lagrangian Theory (q) ^L	-1	$1-q$	0	0	0	0	0	0	0
c. Ni's General Theory (p, q)	-1	$1-q$	0	0	0	0	$4+p-2q$	0	0
d. Nordström	-1	1	0	0	0	0	0	0	0
e. Littlewood-Bergmann	-1	$\frac{1}{2}$	0	0	0	0	-1	0	0
f. Whitrow-Morduch (q)	-1	$1-q$	0	0	0	0	-2q	0	0

TABLE 5.1 (cont'd)

Theory and its adjustable parameters [†]	γ	β	α_1	α_2	α_3	ζ_1	ζ_2	ζ_3	ζ_4
5. Stratified Theories (with Time-Orthogonal Conformally Flat Space Slices)									
a. Rosen (λ) ^L	λ	$\frac{1}{4}(3+\lambda)$	$-4(1+\lambda)$	0	0	0	0	0	0
b. Papapetrou ^L	1	1	-8	-4	0	0	0	0	0
c. Ni's Lagrangian Stratified Theory ^L	1	1	-8	0	0	0	0	0	0
d. Ni's General Stratified Theory (p, q)	1	1-q	-8	0	-4	0	p-2q-2	0	-2
e. Yilmaz	1	1	-8	0	-4	0	-2	0	-2
f. Page-Tupper (a, c)	a	1+c	$-4(1+ca)$	0	$-2(1+ca)$	0	1+a+2c	0	$-(1+ca)$
g. Coleman (p)	1	1	-8	0	-4	0	p-2	0	-2
h. Einstein's 1912 Theory	0	0	-4	0	-2	0	-1	0	-1
i. Whitrow-Morduch Stratified Theory	0	-1	-4	0	-2	0	-3	0	-1
6. Whitehead's Theory	1	1	0	0	0	-6	0	-1	-1
									(plus additional metric terms)

[†] The superscript ^L refers to theories which are Lagrangian-based.

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