MEMORANDUM

November 17, 1987

To: Vanden Bout, Napier, Kellermann From: George Seielstad

Subject: Green Bank fi Environment

- 1) Wes Sizemore has surveyed the rfi as requested. Plots are the worst cases.
- 2) All measurements were made during working day. Times are EST.
- Formulas for converting from dbm (Sizemore's calibration on his spectrum analyzer) to W m⁻² were provided by Fisher.
- 4) Bands chosen for direct comparison with NE results.
- 5) No instrumentation for measurements at frequencies >1500 MHz. Reason is we have essentially no requests for tracking down interference at those frequencies; it simply isn't noticeable.
- 6) Note Sizemore's comments about 300-350 MHz. He has established excellent contacts with the military.

NATIONAL RADIO ASTRONOMY OBSERVATORY Green Bank, WV

MEMORANDUM

November 16, 1987

- To: G. Seielstad
- From: W. Sizemore

Subj: 0-1500 MHz Green Bank RFI Survey

Per your request, the frequency bands

	50- 100	MHz
	73 - 75	MHz
	300- 350	MHz
	500-1000	MHz
	550 - .650	MHz
and	1000-1500	MHz

were surveyed during the week of 9 November 1987. The monitoring procedure, results, and my comments are attached.

WAS/cjd

Enclosure RFI Survey November 9-13, 1987

Monitoring Procedure

In order to conduct this survey, the monitoring station at the 40-ft telescope was used The schematic diagram of this station is attached as Figure 1. The spectrum analyzer was calibrated at 280 MHz using its internal calibration signal a -30 dBm. The following information and formulas were used to convert from the calibrated signal strength shown on the spectrum analyzer to flux density at the antenna.

Conversion to Flux Density

$$S_{(W/m^2)} = \frac{P_a 4\pi}{G_a \lambda_{(m)}^2}$$

where

$$P_{a} = P_{sa}(watts) \left(ALOG \left[\frac{L_{s}(dB) - G_{s}(dB)}{10} \right] \right)$$

$$P_{sa}(watts) = \frac{ALOG \left(\frac{dBm}{10} \right)}{1000}$$

$$\lambda_{(m)} = \frac{300}{f_{(MHz)}}$$

$$G_{a} = ALOG \frac{G_{AD}}{10}$$

where

LS	=	Loss ahead of spectrum analyzer in dB.
G _S	=	Gain ahead of spectrum analyzer in dB.
f	=	Frequency in MHz.
P sa	=	Power at spectrum analyzer input.
dBm	=	Spectrum analyzer reference level.

Results

Plots were taken at azimuths of 0° , 90° , 180° , and 270° for each frequency band. However, I have included only one plot for each frequency band which shows the worst case for that band. These data are presented as Figures 2, 3, and 4.

Comments

- 50-100 MHz: This band contains TV channels 2 to 6, part of FM band, and other fixed mobile transmitters. However, radio astronomy band (73.0 to 74.6 MHz) is fairly quiet.
- 73-75 MHz: The plot of the band contains 3 or 4 "birdies" which may have been due to the close proximity of the monitoring station to the 140-ft and 85-1.
- 300-350 MHz: The predominant user of this band is military aircraft. I have had <u>excellent</u> results in clearing this band for observation by reporting our observing frequencies to the appropriate military command.
- 500-1000 MHz: The majority of this band is allocated to broadcasting which the quiet zone controls very well thank you.
- 550-650 MHz: Allocated to broadcasting except for radio astronomy band. Again, the quiet zone controls this band very well.
- 1000-1500 MHz: The signals centered around 1100 MHz are FAA radars.

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