NATIONAL RADIO ASTRONOMY OBSERVATORY

Charlottesville, Virginia

October 1, 1979 - December 31, 1979

RESEARCH PROGRAMS

Observer(s)

H. Wootten (Caltech)

P. Bowers

M. Reid (Center for Astrophysics)

Hours
1916.00
177.00
51.00
26.50
65.00
7.25
0.00

Program

Studies of the time variations of H_2O

masers in late type stars at 22.4 GHz.

The following line programs were conducted during this quarter.

Observations at 1425 MHz of H166 F. J. Lockman the galactic plane and in the Rosette Nebulae. R. Giovanelli (NAIC, Puerto Rico) Observations of 21-cm hydrogen in sections of the northern Magellanic M. Haynes (NAIC, Puerto Rico) Stream. Search for 21-cm hydrogen in emission B. Balick (Washington) E. Skillman (Washington) from Irr I galaxies. C. Bernes (Stockholm Observatory) Observations of 2-cm H₂CO in the molecular clouds ρ Oph and CrA to use in A. Sandqvist (Stockholm Observaconjunction with 2-mm data to determine tory) R. Loren (Texas) the spatial variation of density. H. Wootten (Caltech) N. Evans (Texas) Studies of the density distribution and R. Loren (Texas) mass spectrum of fragments in selected R. Snell (Texas) molecular clouds by the observation of

2-cm H₂CO.

M. Cohen (Cornell)
B. Turner

P. Bowers

K. Fox (Tennessee)

D. Jennings (NASA, Greenbelt)

Program

Search at 5.625 GHz for ${\rm H_2}^{18}{\rm O}$ maser emission during the "flare" phase of the Ori (KL) ${\rm H_2O}$ maser.

Studies of 18-cm OH emission from unidentified Type II OH/IR Stars.

Observations of interstellar methane and cyanodiacetylene radio emission lines found in the frequency range of 18.5-21.9 GHz.

The following continuum programs were conducted during this quarter.

Observer(s)

Observer(s)

C. Crannell (NASA, Greenbelt)
R. Hobbs (NASA, Greenbelt)

S. Kahler (American Science and Engineering)

W. B. Burton (Minnesota)

H. Liszt

W. B. Burton (Minnesota)

W. Jaffe

H. Liszt

M. Kundu (Maryland)

P. Bowers

Program

Observations at several frequencies between 1-2 GHz of the flare star ZY Canis Minoris, coordinated with other observations.

Observations at 1 GHz to search for evidence of the tilted fundamental inner galaxy gas distribution in thermal and nonthermal radiation.

Observations at 1 GHz of the galactic disk.

Search for 1.3-cm radio emission from cool stars.

The following very-long baseline programs were conducted, and the stations used in the experiments are coded as follows:

A - Algonquin, Canada, 150-ft

B - MPIR, Bonn, FRG, 100-m

C - Chilbolton, England, 25-m

E - Jodrell Bank, England, 250-ft

F - Harvard/Fort Davis, 85-ft

G - NRAO, 140-ft

K - Haystack, 120-ft

0 - OVRO, 130-ft

P - Arecibo, 1000-ft

R - Simeis, USSR, 22-m

S - Onsala, Sweden, 84-ft

- M. Claussen (Iowa)
- J. Fix (Iowa)
- R. Mutel (Iowa)
- J. Benson
- R. Porcas (MPIR, Bonn)
- R. Linfield (Caltech)
- J. Romney (MPIR, Bonn)
- R. Preston (JPL)
- R. C. Walker (Caltech)
- K. Kellermann (MPIR, Bonn)
- B. Geldzahler (Pennsylvania)
- T. Menon (NASA, Greenbelt)
- R. Booth (Jodrell Bank)
- P. Wilkinson (Jodrell Bank)
- A. Readhead (Caltech)
- L. Kogan (Inst. for Space Research, Moscow)
- V. Kostenko (Inst. for Space Research, Moscow)
- L. Matveyenko (Inst. for Space Research, Moscow)
- B. Ronnang (Chalmers, Sweden)
- D. Downes (MPIR, Bonn)
- R. Genzel (Center for Astrophys.)
- J. Moran (Center for Astrophys.)
- M. Reid (Center for Astrophys.)
- K. Kellermann (MPIR, Bonn)
- I. Pauliny-Toth (MPIR, Bonn)
- T. Clark (NASA, Greenbelt)
- R. Coates (NASA, Greenbelt)
- C. Ma (NASA, Greenbelt)
- J. Ryan (NASA, Greenbelt)
- D. Shaffer (Phoenix Corp.)

Program

Observations of 1665, 1667 MHz OH emission from long period variable stars with telescopes F, P, and G_{\bullet}

Observations at 2.8-cm of the quasar 0.723+67 with telescopes B, C, K, O, and G.

Observations at 2.8-cm of the nuclei of radio galaxies with telescopes A, B, F, K, 0, and G.

Observations at 2.8-cm of high redshift QSO's with telescopes B, F, K, O, and G.

Observations at 2.8-cm of the nuclear components of Cyg A and 3C 154 with telescopes at B and G.

Observations at 610 MHz to monitor suspected low-frequency variables and to map selected bright sources with telescopes J, F, O, and G.

Observations at 22 GHz of proper motions in strong $\rm H_2O$ maser sources with telescopes B, R, S, K, O, and G.

Observations at 3.6 and 13-cm to test the capabilities of the Mark III VLBI system and to observe the structure of galactic nuclei, low surface brightness objects, radio stars, and to improve astrometric and geodetic measurements with telescopes B, S, K, O, and G.

Program

(continued)

N. Vandenberg (NASA, Greenbelt)

C. Knight (Haystack)

H. Hinteregger (Haystack)

E. Nesman (Haystack)

A. Rogers (Haystack)

A. Whitney (Haystack)

W. Cotton (MIT)

B. Corey (MIT)

C. Counselman (MIT)

T. Herring (MIT)

I. Shapiro (MIT)

J. Wittels (MIT)

D. Robertson (NGS, Rockville)

R. Lacasse

B. Rayhrer

W. Cotton (MIT)

M. Gorenstein (MIT)

J. Marcaide (MIT)

I. Shapiro (MIT)

R. Preston (JPL)

K. Kellermann (MPIR, Bonn)

A. Witzel (MPIR, Bonn)

M. Reid (Center for Astrophys.)

D. Shaffer (Phoenix Corp.)

F. Owen

M. Davis (Center for Astrophys.)

A. Haschick (Center for Astrophys.)

A. E. Lilley (Harvard)

J. Moran (Center for Astrophys.)

T. Pearson (Caltech)

A. Readhead (Caltech)

M. Cohen (Caltech)

A. Moffet (Caltech)

T. Pearson (Caltech)

A. Readhead (Caltech)

R. Simon (Caltech)

G. Seielstad (Caltech)

R. C. Walker (Caltech)

Observations at 13-cm of the twin quasars 0957+56.1 with telescopes B, K, 0, and G.

Studies at 6-cm of quasar proper motion with telescopes B, K, O, and G.

Observations at 6-cm of the small scale structure of the twin quasar 0947+56.1 with telescopes B, K, O, and G.

Observations at 6-cm to continue the hybrid mapping of a complete sample of sources which may be new superliminal sources with telescopes F, H, O, and G.

Observations to monitor superluminal sources at 6-cm with telescopes B, F, O, and G.

A. Witzel (MPIR, Bonn)

- S. Neff (Virginia)
- K. Johnston (NRL)
- J. Spencer (NRL)
- E. Waltman (NRL)
- R. Brown
- E. Fomalont
- R. Perley
- R. C. Walker

J. Cordes (Massachusetts)

- K. Johnston (NRL)
- T. Jones (Minnesota)
- S. Spangler

Program

Observations at 6-cm to investigate the spatial structure of a complete sample of compact radio sources with telescopes B, F, K, N, O, and G.

Observations at 18-cm to test whether 10 selected sources are inhomogeneous synchrotron sources or sources which produce synchrotron radiation from a relativistic Maxwellian electron distribution, with telescopes F, O, and G.

300-Foot Telescope	Hours
Scheduled observing	1728.50
Scheduled maintenance and equipment changes	136.50
Scheduled tests and calibration	225.00
Time lost due to: equipment failure	27.25
power	52.75
weather	0.00
interference	0.00

The following line programs were conducted during this quarter.

Observer(s)

R. Terlevich (Inst. of Astronomy, Cambridge)

- T. Thuan (Virginia)
- G. Martin (Virginia)
- L. Davis (Toronto)
- E. Seaquist (Toronto)
- D. Burstein
- K. Lo (Caltech)
- C. Kowal (Caltech)
- J. Pier (Caltech)
- W. Sargent (Caltech)

Program

Measurements of the 21-cm hydrogen content of blue compact objects discovered on the 1.2 meter U.K. Schmidt objective prism plates.

Studies of 21-cm hydrogen in isolated and interacting galaxies.

Survey the thinnest galaxies found in the Uppsala General Catalog for 21-cm hydrogen.

Studies of 21-cm hydrogen in dwarf galaxies.

W. Ford (DTM)

V. Rubin (DTM)

N. Thonnard (DTM)

B. M. Lewis (Carter Observatory, New Zealand)

D. Hogg

T. Bania (NAIC, Puerto Rico)

F. J. Lockman

Program

Studies of the extent of 21-cm hydrogen in high luminosity spiral galaxies.

Observations of 21-cm hydrogen in Shapley Ames and other galaxies.

Mapping of 21-cm hydrogen surrounding three Wolf-Rayet N-type stars.

Measurements of the scale height of galactic HI.

The following continuum programs were conducted during this quarter.

Observer(s)

T. Balonek (Massachusetts)

W. Dent (Massachusetts)

C. O'Dea (Massachusetts)

C. Crannell (NASA, Greenbelt)

R. Hobbs (NASA, Greenbelt)

S. Kahler (American Science and Engineering)

J. Cordes (Cornell)

J. Dickey

Program

Polarization and flux density measurements of variable radio sources at 2695 MHz.

Observations at 4 four discrete freqquencies between 250 and 500 MHz of the flare star ZY Canis Minoris coordinated with other observations.

Interstellar scintillation measurements of compact sources at 850 MHz.

The following pulsar programs were conducted during this quarter.

Observer(s)

P. Backus (Massachusetts)

J. Taylor (Massachusetts)

M. Damashek

Program

Observations at 410, 610, and 850 MHz to determine periods, period derivatives, positions and dispersion measures of known pulsars.

J. Taylor (Massachusetts)

M. Damashek

Observations to complete a Northern Hemisphere pulsar search at 410 MHz.

36-Foot Telescope

		Hours	
Scheduled observing 1965.50 Scheduled maintenance and equipment changes 94.50 Scheduled tests and calibration 100.00			
Time not scheduled		48.00	
Time lost due to:	telescope	5.50	
	equipment	45.00	
	weather	53.50	
	power	2.00	
	interference	0.00	

Observer(s)

Program

	Barrett (MIT) Buxton (MIT)	Study of CO in small reflection nebulae.
	B. Burton (Minnesota) Liszt	Study of the thickness of CO in the inner galaxy.
С.	Churchwell (Wisconsin) Winnewisser (MPIR) Walmsley (MPIR)	Observations of methyl formate at 90 GHz.
	Dent (Massachusetts) Hobbs (NASA, Greenbelt)	Evolution of extragalactic radio sources at millimeter wavelengths.
L.	Denoyer (Cornell)	Study of shocked CO within the supernova remnant IC 443.
L.	Denoyer (Cornell)	Search for shock excited molecules near supernova remnants.
D.	Dickinson (JPL)	Observations of chemistry in the regions of interstellar shock waves.
	Jackson (Maryland) Sewall (Maryland)	Study of molecular clouds near HII regions that are distant from the galactic center.
	Jaffe (Center for Astrophys.) Fazio (Center for Astrophys.)	Continuation of CO survey of IR sources in the galaxy.
G.	Jaffe (Center for Astrophys.) Fazio (Center for Astrophys.) McBreen (Center for Astrophys.)	Determination of distance of IR sources by their CO emission.

D. Johnson (Florida)

- S. Gottesman (Florida)
- M. Kutner (Rensselaer)
- N. Evans (Texas)
- K. Tucker (Fordham)
- M. Kutner (Rensselaer)
- N. Evans (Texas)
- K. Tucker (Fordham)
- M. Kutner (Rensselaer)
- N. Evans (Texas)
- K. Tucker (Fordham)
- M. Kutner (Rensselaer)
- K. Tucker (Fordham)
- M. Kutner (Rensselaer)
- D. Machnik (Rensselaer)
- D. Hacimik (Kensselael)
- M. Kutner (Rensselaer)
- D. Machnik (Rensselaer)
- K. Tucker (Fordham)
- C. Leung (Rensselaer)
- R. Levreault (Texas)
- N. Evans (Texas)
- R. Loren
- K. Mattilla (Helsinki Observatory)
- L. Malkamaki (Helsinki Observatory)
- G. Snadell (Chalmers)
- M. Morris (Columbia)
- B. Elmegreen (Columbia)
- D. Elmegreen (Hale Observatories)
- P. Myers (MIT)
- R. Buxton (MIT)
- R. Plambeck (Calif., Berkeley)
- R. Loren (Texas)
- R. Snell
- L. Rickard
- P. Palmer

Program

Confirmation of the detection of CO in NGC 185.

Study of formaldehyde isotopic abundances at 2 millimeters.

Testing fractionation effects in NGC 1977 using DCO^+ and $^{13}CO^+$.

Study of H_2CO and $H_2^{13}CO$.

Study of methane isotopes at 2 milli-meters.

High resolution maps of dsense molecular clouds in Canis Majoris.

Studies of broad CO lines in T Tauri stars.

Study of CO line profiles toward FU Ori, variable stars and other objects.

Mesurements of CO and ¹³CO in five dark nebulae.

Study of CO in selected irregular galaxies.

Study of gas motions in dark clouds by CO emission.

Mesurement of CO 1-2 for comparison with measurements at Texas of the 2-1 CO lines.

Continuation of the search for extended CO emission in spiral galaxies.

Program

L. Rudnick (Minnesota) T. Jones (Minnesota) R. Fiedler (Minnesota) J. Puschell (Minnesota)	Polarization observations of extragalactic radio sources.
P. Schwartz (NRL)	Observations of thermal emission from dust at 2 millimeters.
B. Wilking (Arizona) C. Lada (Arizona)	Study of self reversed CO lines from dark clouds.
E. Wright (MIT)	Maps of CO emission from far IR sources.

The Very Large Array

B. Zuckerman (Maryland)

The array was scheduled for observations 55 percent of the time in the fourth quarter of 1979. Forty-one percent of the time was devoted to astronomical observing and the remaining 14 percent to instrumental development and tests. Approximately 19 percent of the observing time was lost to instrumental problems. The following research programs were conducted with the VLA during this quarter.

Observer(s)

Program

Search for CO emission from stars in the

IR 2 micron catalogue.

B. Balick (Washington) P. Crane T. Heckman (NFRA, Leiden)	Synthesis of Seyfert galaxies. 6 and 20 cm.
R. Becker (NASA, Greenbelt)	SNR G21.5-09. 6 and 21 cm.
R. C. Bignell E. R. Seaquist (Toronto)	SNR in NGC 4449. 2, 6, and 21 cm
A. Bridle (Queen's) E. Fomalont J. Hogbom (Stockholm Observatory) A. Willis (NFRA, Zwiggelte)	3C 315. 2 cm.
B. Burke (MIT) D. Roberts (MIT)	Double quasar 0957+561. 2, 6, and 20 cm.
J. Burns S. Gregory (Bowling Green SU)	4C sources in poor Zwicky clusters. 20 cm.

H. Johnson (Lockheed)
W. Cash (Calif., Berkeley)

Program

Radio star astrometry. 6 cm.

J. Condon J. Mitchell (Penn State) P. Usher (Penn State)	Optically selected QSO's; K optical variables. 6 cm.
M. Felli (Arcetri)N. Panagia (Laboratorio di Radio Astronomia, Bologna)	Massive stars undergoing mass loss. 6 cm.
E. Fomalont A. Bridle (Queen's)	X-ray cluster Abell 2256. 6 and 20 cm.
B. Geldzahler (MIT) E. Fomalont	Central region of Fornax A. 6 and 20 cm.
R. Genzel (Center for Astrophys.) J. Moran (Center for Astrophys.) D. Downes (MPIR) M. Reid (Center for Astrophys.)	Very compact HII regions towards OH/H ₂ O maser sources. 6 and 2 cm.
D. Gibson (New Mexico Tech)	Antares. 6 and 20 cm.
A. Haschick (Center for Astrophys.) P. Crane J. van der Hulst (Minnesota)	Neutral hydrogen absorption in NGC 1275. 21 cm line.
D. Heeschen J. Wrobel (Toronto)	Active E/SO galaxies. 2, 6 and 21 cm
D. Heeschen D. Hogg H. Arp (Hale Observatories)	Maps of fields around seven interesting galaxies. 6 and 21 cm.
R. Hjellming G. Miley (NFRA, Leiden)	Moving jets in SS 433. 1.3, 2, 6, and 20 cm.
P. Ho (Calif., Berkeley) A. Haschick (SAO)	Low mass star formation near HII regions. 6 and 20 cm.
W. Jaffe C. Perola (Milano, Italy) G. Gavazzi (Milano, Italy)	Bright spiral galaxies in nearby rich clusters. 21 cm.

Program

K. Johnston (NRL)C. WadeD. Gibson	Radio star astrometry. 6 cm.
K. Johnston (NRL)S. Wyckoff (Arizona State)L. Rudnick (Minnesota)F. Ghigo (Minnesota)	Comparison of optical and radio structure of 3 low-redshift QSO's0837-12, 0812+020 and 0736+014. 2, 6, and 20 cm.
P. Kronberg (Toronto) P. Biermann (MPIR)	High-z QSO's 0225-014 and 0642+44. 6 and 20 cm.
M. Kundu (Maryland) A. Rao (Maryland)	Solar active regionssimultaneous radio and x-ray observations. 6 and 21 cm.
M. Kundu (Maryland)E. Schmahl (Maryland)T. Velusamy (Maryland)	Gradual and rapid evolution of solar active regions. 6 and 20 cm.
<pre>K. Lang (Tufts) R. Willson (Tufts)</pre>	Solar active regions. 2, 6, and 20 cm.
P. Lindblad (ESO, Switzerland) A. Sandqvist (Stockholm Observtory) S. Jorsater (Stockholm Observatory)	Continuum study of barred galaxy NGC 1365. 6 and 20 cm.
<pre>K. Lo (Caltech) R. Brown K. Johnston (NRL)</pre>	Galactic center, the compact object. 6 and 20 cm.
C. Masson (Caltech)	Clusters of galaxies with complicated 151 MHz structure. 6 and 21 cm.
J. Moran (Center for Astrophys.) L. Rodriguez (Mexico)	NGC 6334S. 6 and 20 cm.
S. Neff (Virginia) R. Brown	Optical spectra vs radio structure of QSO's. 6 and 18 cm.
R. Newell (New Mexico Tech) R. Hjellming	Mapping of thermal star shells. 2 and 6 cm.
R. Perley A. Willis (NFRA, Zwiggelte)	NGC 6251. 20 cm.

Observer(s) Program Halo of Perseus A. 20 cm. R. Perley K. Johnston (NRL) P. Crane R. Potash (Brandeis) 4C 32.69--a quasar with a beam. 2 cm. J. Wardle A. Readhead (Caltech) 3C 48, 3C 138, 3C 147, 3C 245, 3C 309.1, P. Wilkinson (Jodrell Bank) 3C 380, and 3C 454.3. 1.3, 2, and 6 cm. P. Napier R. C. Bignell J. Roberts (Calif., Berkeley) Jovian Radiation Belts. 21 cm. R. C. Bignell G. Berge (Caltech) E. Seaquist (Toronto) SS 433. 1.3, 2, 6, and 20 cm. W. Gilmore (Toronto) D. Shaffer (Phoenix Corp.) Optically discovered quasars. 6 cm. R. Green (Caltech) M. Schmidt (Caltech) R. Sinha NGC 6217, a barred spiral. 20 cm. V. Rubin (DTM) J. Vallee (Queen's) X-ray galaxy NGC 3862 = 3C 264. 20 cm. A. Bridle (Queen's) J. van der Hulst (Minnesota) Stephan's Quintet. 21 cm. A. Rots

ELECTRONICS DIVISION

P. Freier (Minnesota)

Charlottesville

J. van der Hulst (Minnesota)

J. Young (Massachusetts)

Work has commenced on millimeter-wave frequency doublers and one unit providing a minimum 2.5 mW from 140 to 150 GHz has been shipped to Tucson. A model of the multiplier operating at an input frequency of approximately 2 GHz has been constructed and is being evaluated to gain an understanding of the device. The goal of this project is to provide doublers giving sufficient power (>2 mW) for local oscillator use with cooled mixers in the 115-260 GHz range.

z-distribution of nonthermal continuum

radiation of edge-on galaxies.

Work continues on both construction and development of cryogenically-cooled GASFET amplifiers in the 5 GHz range. Several units have been shipped to the VLA for second-stage use and more units are under construction for use at the VLA, Tucson, Green Bank, and Fort Davis, Texas. A noise temperature of 19 K, including an input isolator, has been achieved in a two-stage amplifier.

Investigation of a mathematical analysis method for evaluation of mixers by back-short vs. diode current data is continuing along with the investigation of improved mixers for the 70 to 115 GHz band.

Expansion of the VLBI Mark II processor and construction of the Model IV autocorrelator are continuing.

Green Bank

The 5-25 GHz upconverter/maser receiver for the 140-foot telescope is nearing completion. The dewar and cryogenic components have been assembled and cryogenic tests are in progress to determine the thermal load on the refrigerator and cool down time. The installation at the Cassegrain focus is planned for summer 1980.

The 300-1000 cooled upconverter/GASFET amplifier receiver for the 300-foot telescope is progressing and should be ready for use on the 300-foot telescope during the summer of this year.

The development of a maser for $40-50~\mathrm{GHz}$ is continuing. A more suitable superconducting magnet has been designed and is being manufactured.

The 22 GHz focal plane measurement receiver is now almost assembled and will be evaluated on the 140-foot telecope for antenna surface measurements.

The IF section of the Model IV correlator is now complete and will be tested in conjunction with the digital section during February in Charlottesville. The current plan is to move both sections to Green Bank during March.

Tucson Electronics

New diodes for the 70-115 GHz band have been installed during this quarter. The receiver noise temperature over the band varies from 284 K SSB to 450 K SSB. We believe that the 284 K is the lowest receiver temperature obtained for a resistive mixer at $100~\mathrm{GHz}$.

During this quarter a path length modulator has been added to the system to suppress the standing waves that are troublesome in spectral line work. As a result we are able to frequency switch for small (up to 10 MHz) excursions in frequency. Residual baseline effects due to the receiver remain when frequency switching, and we hope to cure these in the near future.

The new cooled 2-mm receiver was tested on the telescope during this quarter. The noise temperature was $800~\rm K$ SSB and the stability both for continuum and spectral line work was good. We now have a doubler that will give us coverage from $140-155~\rm GHz$, and this will enable us to abandon klystrons in this frequency range.

The He³ bolometer system has been delayed owing to a series of small vacuum leaks. The quasi optical bandpass filters that are essential for the optimum performance of this receiver are finished and have been tested. We have filters for the 1 mm, 2 mm and 3 mm atmospheric windows. The transmission of the filters varies from 0.9 for the 3 mm to 0.65 for the 1 mm. A new beam switching mechanism has been fabricated for this receiver and will shortly be tested.

ENGINEERING DIVISION

Design of a new spoiler reflector for the 140-foot Cassegrain house was completed and drawings sent to the shop for fabrication. Drawings for the 300-foot traveling feed system were completed and sent to the shop for fabrication. Additional checks and studies were made on the springs for the 140-foot brakes. Modifications to the deformable subreflector were completed and the unit was put in operation. The prototye cast aluminum surface plate was re-measured and the results evaluated. Assistance was provided in negotiating a contract for a new cover for the proposed 25-meter millimeter wave telescope. Routine engineering assistance was provided to operations and maintenance in Charlottesville, Green Bank, Tucson, and the VLA.

COMPUTER DIVISION

IBM 360/65

A new Memorex disk controller has arrived and will be installed soon. This will connect two I/O channels to four disks each rather than the present system of one channel to eight disks, plus another channel for error re-try.

VLA Post Processing

The Digital Equipment Corporation VAX 11/780 digital computer has been installed at the Edgemont Road building in Charlottesville and has been accepted. Post processing programs for the VLA and VLBI are being developed.

The post processing group's contingent of eight persons is complete, with people already at Charlottesville or offers accepted.

Green Bank

A "Pandora" terminal has been installed at Green Bank, connecting users there directly to the Pandora text editing facility of Charlottesville's IBM 360 computer. A Mod Comp Model II/25 digital computer with 64k 16 bit words of storage has been installed at the Green Bank lab building. This computer will be available to telescope users for further data analysis.

VERY LARGE ARRAY

The array was scheduled for oservations and tests for approximately 55 percent of the time during the third quarter. The maximum number of antennas used for an astronomical observing program during the month of December was 19, increasing the number of interferometer baselines to 171. To date, 23 antennas have obtained first fringes. Our longest astronomically usable baseline is approximately 24 km. During the quarter the first VLA scientific spectral line research observation was completed.

On November 9, 1979, Antenna No. 28 was accepted from E-Systems. The completion of this phase of the construction was three weeks ahead of the schedule established in July 1976, and 16 months ahead of the original subcontract schedule.

Retrofitting of the new spectral line baseband systems was completed on the East arm and all but one rack on the North arm. The first retrofit of the new front end IF subsystem was started on Antenna 13 and will be implemented on all other antennas in 1980.

Procurement of additional terminals was started, for use with the Modcomp on-line systems in implementation of the final control center/ maintenance center display system. The CLEAN algorithm, as implemented on the PDP-11/70, is proving a very successful system and is in heavy use. Procurement process is in the first stages for tape drives for the PDP-11 computers.

By the end of the fourth quarter of 1979, Phase IV Site & Wye construction work was 97 percent complete, and Phave V Site and Wye work was 75 percent complete.

The condemnation lawsuit for land acquisition was finally settled on November 20th by the payment of \$205,400 to the Registry of the U.S. Court.

PERSONNEL

Appointments

Jeffrey J. Puschell	Research Associate	10-01-79
Roger D. Blandford	Visiting Scientist	10-03-79
John M. Dickey	Assistant Scientist	10-15-79
Galen R. Gisler	Assistant Scientist	11-01-79
R. Craig Walker	Assistant Scientist	11-13-79
Terminations		
Malcolm W. Sinclair	Electronics Engineer I	10-05-79
Bobby L. Ulich	Electronics Engineer I/	
•	Head, Telescope Operations Division	12-07-79
Roger D. Blandford	Visiting Scientist	12-18-79

APPENDIX A

A list of Observatory reprints issued since January 1, 1979.

Series A

No.	Title	Author(s)	Journal
907	Isotope Abundance Anomalies in IRC +10216	P.G. Wannier R.A. Linke	Astrophys. J., 225, 130-137, 1978.
908	Neutral Hydrogen Absorption Within the Radio Galaxy 3C 178	A.D. Haschick W.A. Baan B.F. Burke	Astrophys. J., 225, 343-345, 1978.
909	Effects of Expanding Compact H II Regions upon Molecular Clouds: Molecular Dissociation Waves, Shock Waves, and Carbon Ionization	J.K. Hill D.J. Hollenbach	Astrophys. J., 225, 390-404, 1978.
910	Radiative-Transfer Effects and the Interpretation of Inter- stellar Molecular Cloud Obser- vations. I. Basic Physics of Line Formation	C.M. Leung	Astrophys. J., 225, 427-441, 1978.
911	VLBI Observations of III Zw 2	J.J. Wittels W.D. Cotton I.I. Shapiro	Astrophys. J., 225, L47-L48, 1978.
912	Observations of Strongly Deuterated Molecules: Impli- cations for Interstellar Chemistry	B.E. Turner B. Zuckerman	Astrophys. J., 225, L75-L79, 1978.
913	Star Formation in the Bright- Rimmed Molecular Cloud IC 1848 A	R.B. Loren H.A. Wootten	Astrophys. J., 225, L81-L84, 1978.
914	A Search for HDCO in Dark Clouds	P.E. Angerhofer G.S. Rossano W.T. Vestrand	Astron. J., 83, 1417 1419, 1978.
915	Radio Sources in Globular Cluster Fields	R.T. Rood K.C. Turner S.J. Goldstein	Astrophys. J., 225, 804-807, 1978.
916	Molecular Hydrogen in Globular Clusters: A Search for Carbon Monoxide	M.H. Schneps P.T.P. Ho A.H. Barrett R.B. Buxton P.C. Myers	Astrophys. J., 225, 808-814, 1978.

No.	Title	Author(s)	Journal
917	The Gas Distribution in the Central Region of the Galaxy. I. Atomic Hydrogen	W.B. Burton H.S. Liszt	Astrophys. J., 225, 815-842, 1978.
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No.	Title	Author(s)	Journal
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Series B

No.	Title	Author(s)	Journal
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